Beyond the Standard Model

The Standard Model

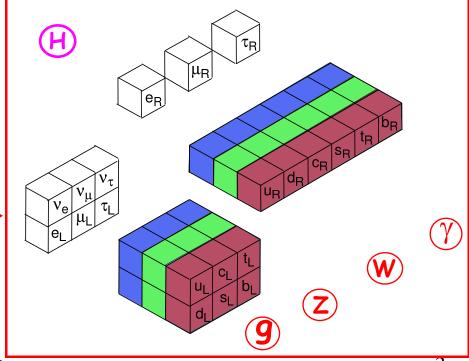


- lacktriangle The standard model \Longrightarrow the electromagnetic, strong and weak interactions \Longrightarrow based on the principle of gauge invariance.
- Lots of free parameters:
 quark and neutrino mixing parameters,

lepton and quark masses, coupling constants,

W, Z and H masses...

Basic fermions and gauge bosons

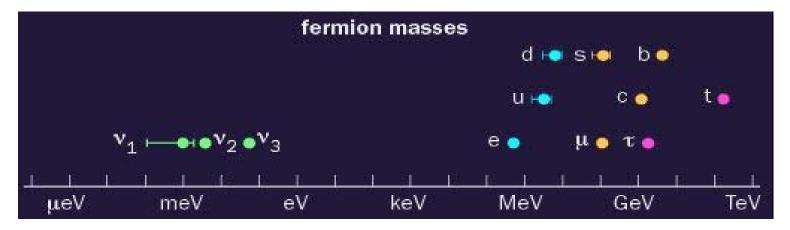


The Standard Model

- The Standard Model agrees very well with all experimental data.
- \bullet The model has been tested down to 10^{-18} m.
- It has been tested to a precision better than 0.1%.
 - Problems with the standard model:
- Does the Higgs exist?
- If neutrinos have mass, are there right-handed neutrinos?
- Why are there 3 generations?
- What about gravity?

The Standard Model

Why are the masses so different (the hierarchy problem)?

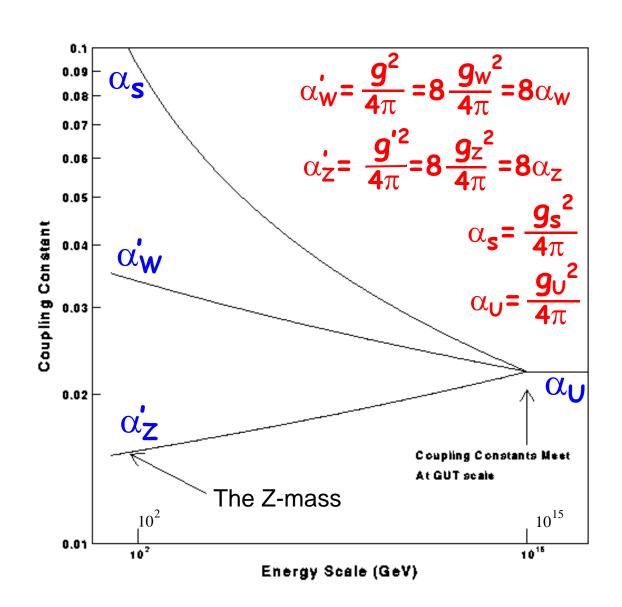


- Can the strong and electroweak interaction be described by a unified theory?
- What happened with the anti-matter in the Big Bang?
- What is dark matter?
- What is dark energy?



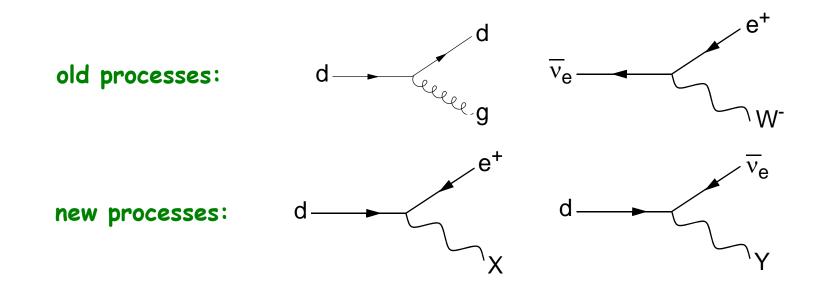
The Georgi-Glashow model

- Weak and electromagnetic interactions are unified.
 - > add the strong one!
- Coupling constants are not truely constant \Longrightarrow they depend on energy (or Q²) in the interaction.
- Unification at some very high unification mass \Longrightarrow electroweak and strong couplings become equal.



- Grand unified theories \implies constructed in many different ways.
- Example: The Georgi-Glashow model \Longrightarrow combines coloured quarks and leptons in single families \Longrightarrow $(d_r, d_a, d_b, e^+, v_e)$
- Two new gauge bosons are introduced \Longrightarrow X with Q=-4/3 and Y with Q=-1/3
- The gauge bosons have a mass close to the unification energy \implies Extremly heavy: $M_X=10^{15}$ GeV/c²
- Single unified coupling constant (g_U) :

New gauge bosons processes possible in which quarks are transformed into leptons by exchanging X and Y bosons:



• The model predicts a value of α_U & relationship between g_u , g and $g' \Longrightarrow$ it predicts a value for the weak mixing angle:

 $\sin^2 \theta_W = 0.21$ Close to the measured value!

• Prediction \implies the sum of the charges is zero within a family $(d_r, d_g, d_b, e^+, \overline{v}_e)$

The number of colours
$$(d_r, d_g, d_b, e^+, \overline{v}_e)$$

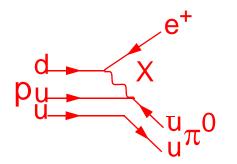
$$3Q_d + e = 0$$

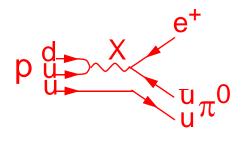
Since d-quark have the charge -e/3 the model works!

- Baryon and lepton numbers are not necessarily conserved in GUT.
- Non-conservation
 —> why the world is dominated by baryons even if the same amount of baryons and anti-baryons were produced in the Big Bang.

Proton decay experiments

• Grand Unified Theories => The proton must be unstable!
It decays by processes involving the X and Y bosons:





Baryon and lepton numbers are not conserved in these processes but the following combination is:

$$B - L \equiv B - \sum_{\alpha} L_{\alpha} \qquad (\alpha = e, \mu, \tau)$$

• It is possible to estimate the lifetime of the proton (τ_p) from a simple zero-range approximation:

$$\tau_p = 10^{32} - 10^{33} \text{ years}$$
 (Age of universe = 10¹⁰ years)

Proton decay experiments

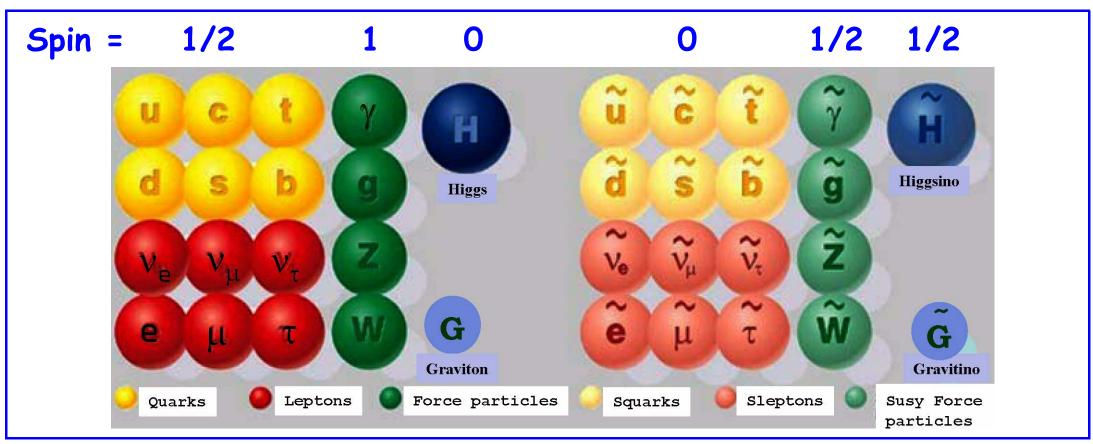
- Many experiments that are doing neutrino physics (Kamiokande, IMB) started out as proton decay experiments.
- The most searched for decay mode $\implies p \to \pi^0 + e^+ \to \gamma \gamma + e^+$ Look for one positron + two electron-positron pairs (from photon conversions).
- No proton decays have been observed \(\subset\) upper limit:

$$\frac{\text{proton lifetime}}{\text{branching ratio}} = \frac{\tau_p}{B(p \to \pi^0 e)} > 5 \times 10^{32} years$$

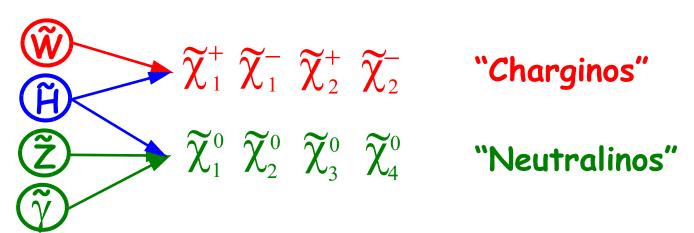
The Georgi-Glashow model predicts:

$$\frac{\text{proton lifetime}}{\text{branching ratio}} = 0.003-0.030 \times 10^{32} \text{ years}$$

- Supersymmetry (SUSY) \Longrightarrow a GUT in which interactions are symmetric under the transformation of a fermion to a boson.
- Every known elementary particle have a super-symmetric partner (superpartner) with different spin.



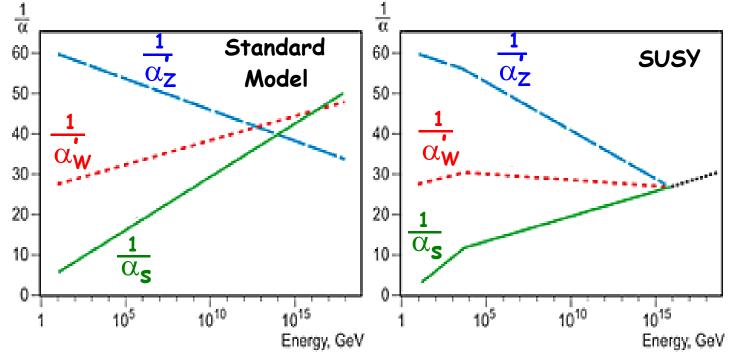
- The new particles are called squarks, sleptons, photinos, gluinos, winos, zinos a tilde is used to denote these particle: \hat{e} \hat{W} $\hat{\gamma}$
- New particles must be heavier than the known particles —>
 otherwise they would already have been discovered.
- The Lightest Supersymmetric Particle (LSP) is stable (in most SUSY models).
- New states are predicted due to mixing between some of the super partner states:



• There are many different supersymmetric models:

<u>Name</u>	<u>LSP</u>	New parameters
Minimal MSSM: super symmetric standard model	Any	> 100
cMSSM: Constrained MSSM	$\mathbf{\tilde{\chi}}_{1}^{0}$	M_0 , $M_{1/2}$, A_0 , $tan(\beta)$, $sgn(\mu)$
mSUGRA: Minimal Supergravity	$\mathbf{\tilde{\chi}}_{1}^{0}$	M_0 , $M_{1/2}$, A_0 , $tan(\beta)$, $sgn(\mu)$
AMSB: Anomaly mediated symmetry breaking	$\mathbf{\tilde{\chi}}_{1}^{0}$	m_0 , $M_{3/2}$, $tan(\beta)$, $sgn(\mu)$
GMSB: Gauge mediated symmetry breaking	Ĝ	$Λ_{m}$, M_{m} , tan(β), N_{5} , sgn(μ)

• Extrapolation of α to the unification scale works better with SUSY.



- SUSY predicts a value for the weak mixing angle which is closer to the experimental results than the Georgi-Glashow model.
- Some SUSY models unify ALL forces including gravity at the Planck mass of 10¹⁹ GeV by replacing particles with superstrings.

- SUSY search in the DELPHI experiment
- SUSY at LEP => Example: selectron production => decay to electrons and neutralinos:

$$e^{+} + e^{-} \rightarrow \tilde{e}^{+} + \tilde{e}^{-}$$

$$e^{+} + \tilde{\chi}_{1}^{0}$$

$$e^{+} + \tilde{\chi}_{1}^{0}$$

- 1) The cross section for producing selectron pairs is comparable to that of producing ordinary charged particles with the same mass.
- 2) The selectrons decay before they can reach a detector.
- 3) The neutralinos only interact weakly and they are therefore virtually undetectable.

- Signature

 > events with only an electron and a positron.
- Selection:
 - i) the e⁺ and e⁻ should carry only half of the collision energy
 - ii) the e⁺ and e⁻ are not emitted in the opposite directions
- No events were found with a clear and background free signature.

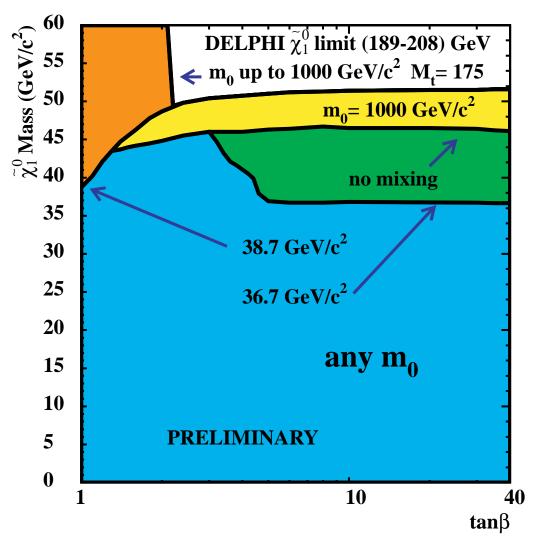
- lacktriangle The results $\locate{}$ set lower limits on the mass of the neutralino.
- SUSY has many models, each with different sets of unknown parameters > the results are given for different assumptions on models and parameters.

- The slepton searches were combined with other SUSY searches to set limits on the neutralino mass.
- All the coloured areas are excluded by the searches.

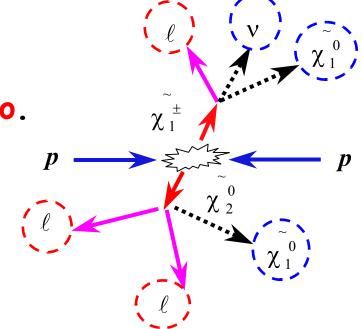
- The tan(β) parameter is related to the SUSY Higgs particles and m₀ to the sfermions mass.
- Result:

 $\tilde{\chi}_{1}^{0}$ > 36.7 GeV

for all parameter values.



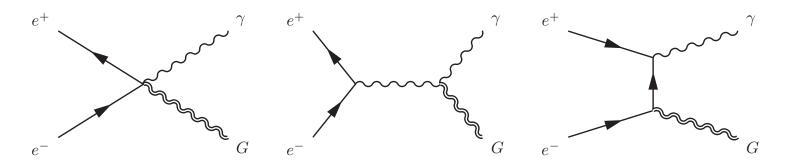
- SUSY search in the ATLAS experiment
- One of the main purposes of the LHC experiments are to search for SUSY as predicted by different models.
- Promising channel:
 the production of a chargino & neutralino
 decays to leptons & lightest neutralino.
- Signature: three leptons (electrons & muons) and missing energy.



- ullet Selection: $P_T > 10$ GeV for the lepton & Missing energy > 30 GeV.
- Background: ZW and tt production.

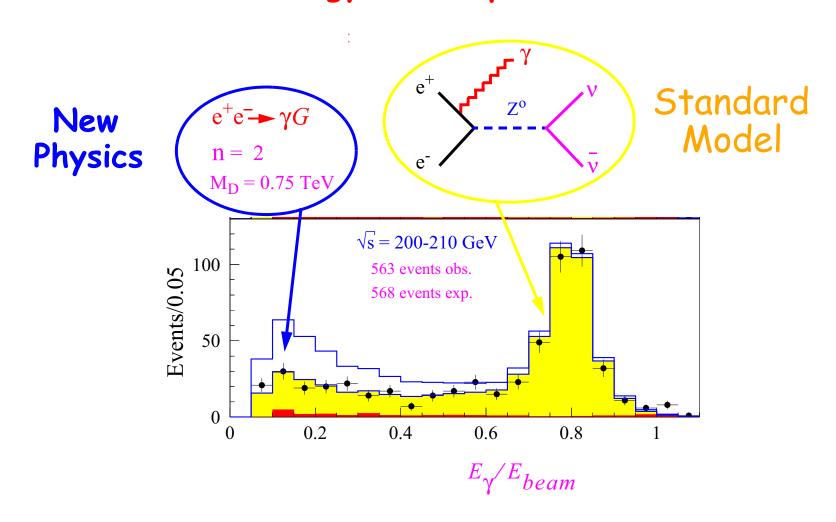
- ullet The gravitational force is much weaker than the electromagnetic and strong interactions \Longrightarrow not studied in particle physics.
- Postulated the existance of graviational force carriers \square \triangleright Gravitons (G).
- Graviation has only been studied at large distances (>1 mm)
 it could be stronger at shorter distances.
- New theories Gravity is unified with other interactions New dimensions of space in which only gravity can propagate (new dimensions in addition to the normal 3 space + 1 time dimensions).

- Graviton search in the DELPHI experiment
- At the energy scale where gravity is unified with other forces
 gravitons are produced in high-energy collisions
 but escape undetected into the extra dimensions.
- Prediction: e⁺e⁻ colliders with sufficient energy \Longrightarrow events with one graviton and one photon:



Gravitons cannot be detected
 only one photon in the experiment.

Selections: Events with only one photon and nothing else.
 Measurement: The energy of the photon



Conclusion: No sign of graviton production in the data!

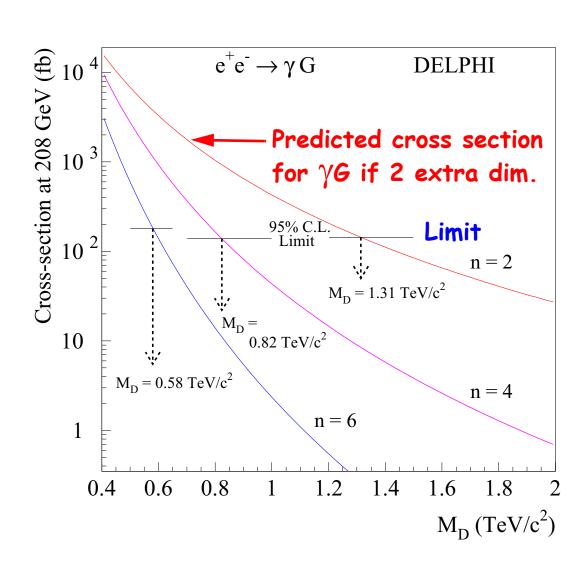
lacktriangle The measurement \Longrightarrow set limits on the parameters in the theory

- Parameters:
 - i) n the number of extra dimensions
 - ii) M_D a fundamental mass scale.
- Limits:

The data set limits on the cross section

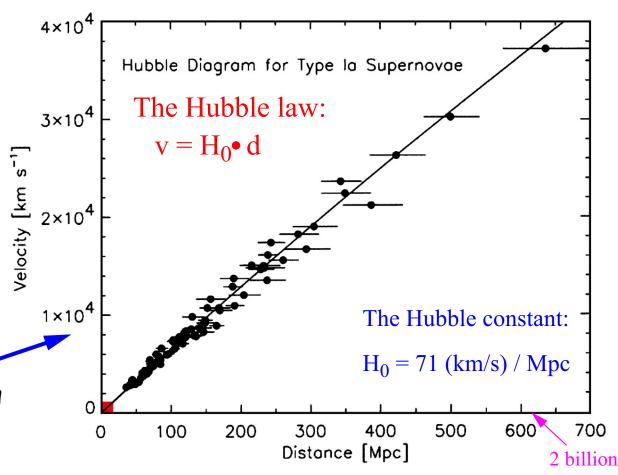


limits on M_D and n in the theory



The Big Bang Model

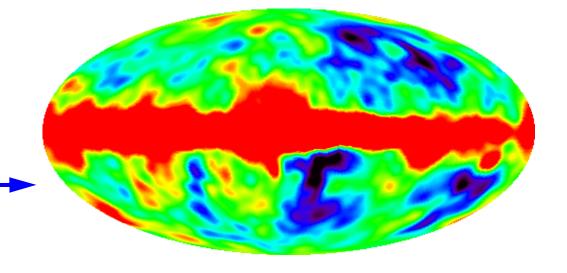
- Experimental evidence for the Big Bang model —>
- 1) A nearly uniform distribution of matter in the universe.
- 2) An abundance of light elements such as He, D and Li.
- 3) The universe is expanding and the velocity of super-novas are therefore increasing with their distance to earth.



4) The cosmic background radiation.

lightyears

The sky as seen at microwave frequencies by the COBE satellite:



The difference between the hottest regions in red and the coldest in blue are only 0.0002 K while the average temperature is 2.7 K.

Conclusion: The cosmic microwave background radiation is very uniform.

The critical density:

Hubble constant $c_c = \frac{3H_0^2}{8\pi G} = O(10^{-26}) \text{ kg m}^3$ Gravitational constant

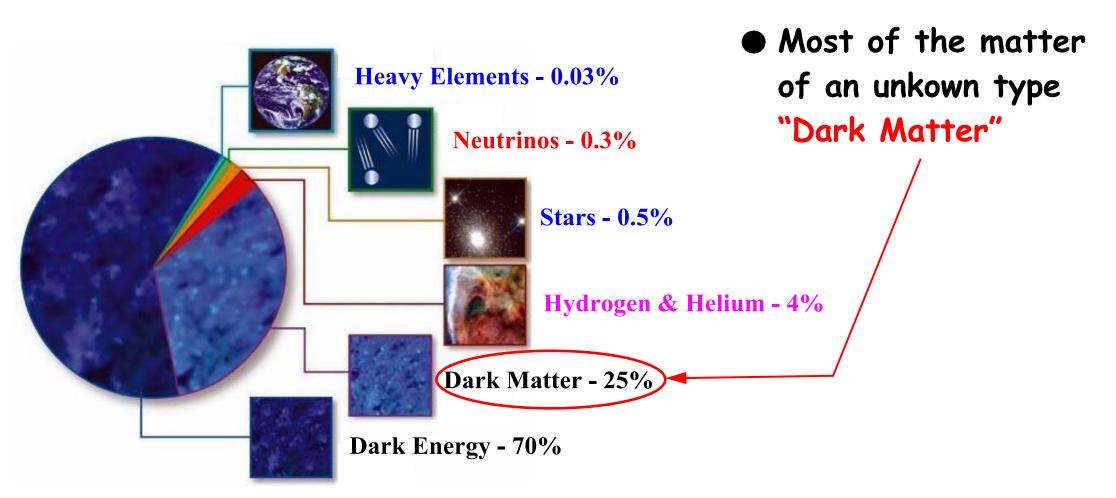
■ If the density in the universe is larger than the critical density
□> the expansion of the universe will eventually end.
(otherwise it will continue for ever).

The relative density:

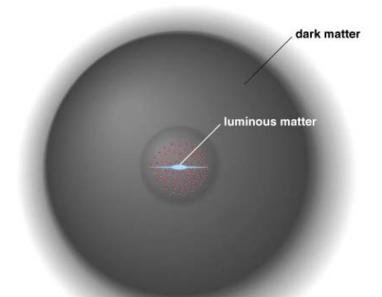
$$\Omega \equiv \rho/\rho_{c} = \Omega_{M} + \Omega_{\Lambda}$$
 Matter part Energy part

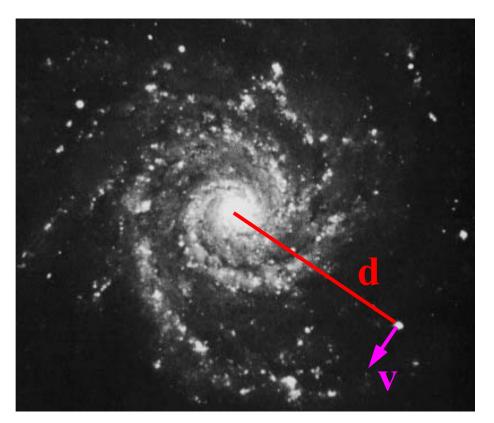
→ Dark Matter

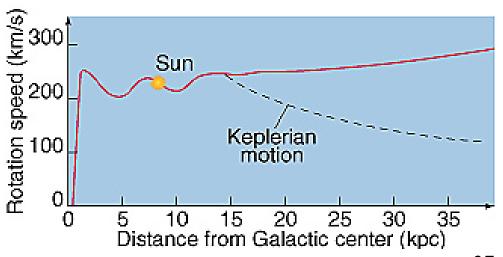
The inflationary Big Bang model =>
the density is close to the critical density.



- Evidence for dark matter measurements of the rotation velocity of stars in galaxies.
- The large rotational velocity of stars in the outer regions of the the Milky way —> the galaxy is full of dark matter.



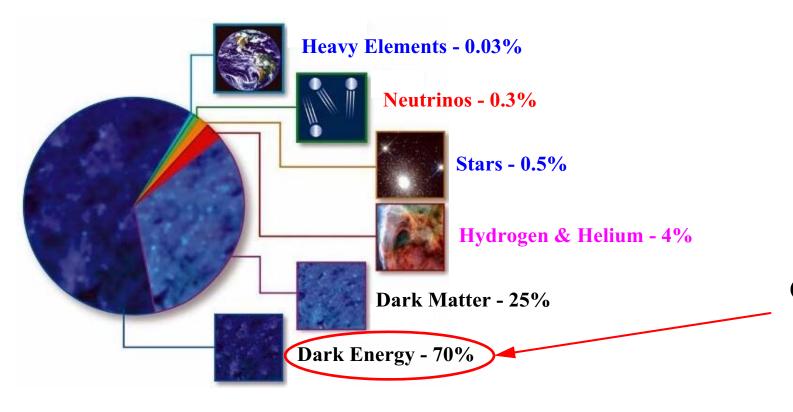




- What does dark matter consist of ?
- 1) Baryonic matter that emits little or no electromagnetic radiation: Brown dwarfs, small black holes MACHO's (for Massive Compact Halo Object).
- 2) Hot dark matter: If neutrinos have a mass > 1 eV \implies give a significant contribution to the density of the universe. But it is difficult to explain how the galaxies are formed if neutrinos make up the dark matter.
- 3) Cold dark matter: Weakly Interacting Massive Particles (WIMPs). Non-baryonic objects that were non-relativistic at the early stages of the evolution of the universe.

 SUSY particles could be WIMPs.

→ Dark Energy

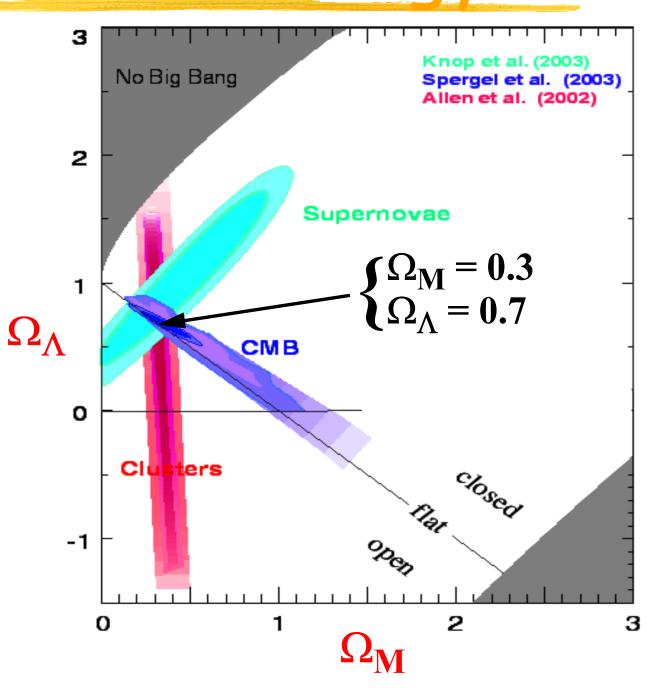


Dark Matter: gravitationally attractive force

Dark Energy: gravitationally repulsive force

Other evidence for dark energy

 studies of the Cosmic Microwave Background (CMB) & the motion of clusters in galaxies.



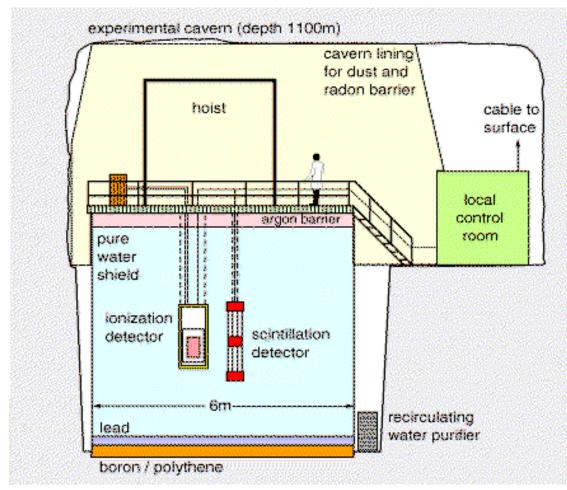
What is causing the Dark Energy?

Two main hypothesis:

- 1) The Cosmological Constant: Space has an intrinsic constant fundamental energy (10⁻²⁹ g/cm³). Calculations of vacuum fluctuations in particle physics give rise to an energy density in vacuum but the calculated value do not agree with astronomical observations.
- 2) Quintessence: Particle-like excitations in a new dynamical field called quintessence. This field differs from the Cosmological constant in that it can vary in space and time.

Direct search for WIMPs

- Interactions between WIMPs and matter has to be very rare.
 □ about one WIMP interacting in a kg of matter every day.
- WIMP detectors are installed deep underground and surrounded with shielding in order to minimize the background.
- The Boulby experiment:
 NaI detector which produces scintillation light if a WIMP interacts with an atom.



200 tons of ultra pure water is used for shielding.

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- The Cryogenic Dark Matter Search (CDMSII)
- In 2009, CDMSII claimed "a hint" of a dark matter discovery.
- CDMSII: Ge detectors at the Soudan underground laboratory to look for WIMPs.
- Interactions between WIMPs and the Ge atoms ⇒ phonons and ionization ⇒ detected by sensors on the semiconductors.
- Two candidate events with 0.9 expected from background.



At a Mine's Bottom, Hints of Dark Matter

By DENNIS OVERBYE

Published: December 17, 2009

An international team of physicists working in the bottom of an old iron mine in Minnesota said Thursday that they might have registered the first faint hints of a ghostly sea of subatomic particles known as dark matter long thought to permeate the cosmos.

