

FYST17 Lecture 11

BSM I

Thanks to G. Broijmans, C. Grojean

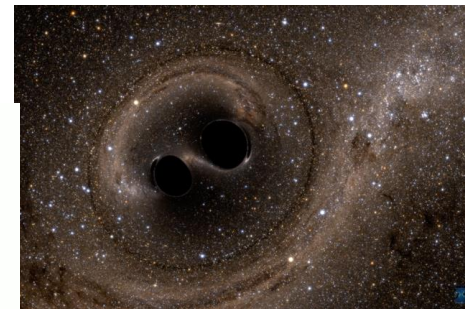
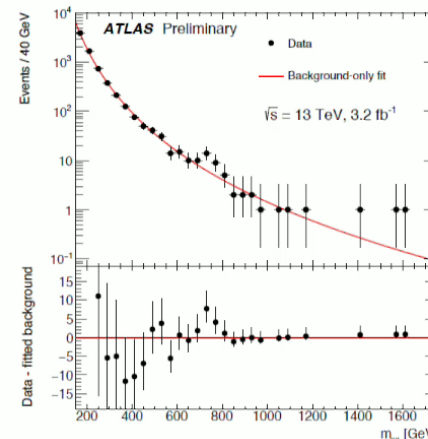
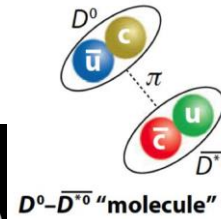
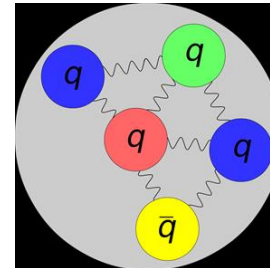
This week's topics

- Why go Beyond the SM?
 - What are the problems with the SM?
 - What direct measurements points to physics BSM
- Some attempts at solutions
 - Supersymmetry
 - Extended Higgs sector
 - Extra dimensions
 - A few others
- Searches for DM

Any direct evidence?

Certainly a few measurements that are not incorporated in the current Standard Model:

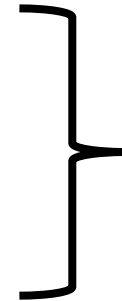
- Exotic baryons (X, pentaquarks etc)
- Neutrino masses!
- (Gravitational waves)
- The new $\gamma\gamma$ bump, if it is real



Status of the Standard Model

19 parameters (+ ν masses)

Tested to precision level $10^{-3} - 10^{-12}$



Extremely
successful!

But empirically incomplete

Structure quite complicated

Aesthetically unacceptable

Many problems with naturalness

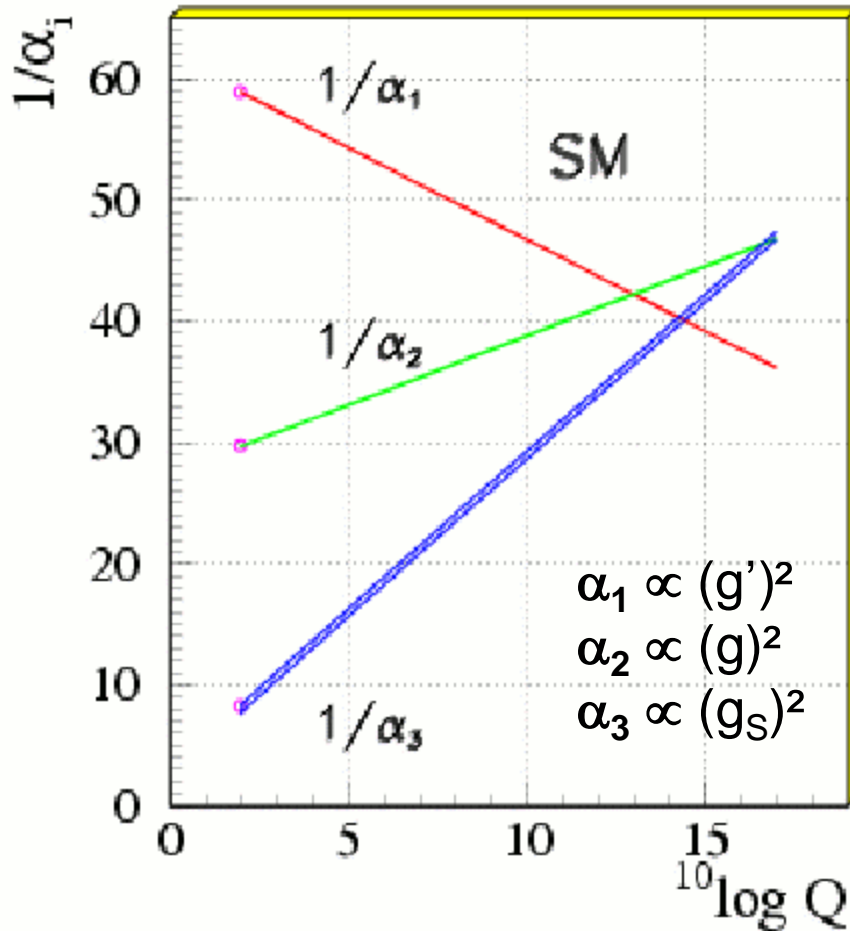
No quantum gravity

Missing answers to "big" questions

Examples of answers we need

- *What is the origin of CP violation?*
- *What is the origin of the matter/anti-matter asymmetry*
- *Why three gauge forces (so far)? And three generations?*
- *Why is the strong interaction strong? Why only left-handed particles participate in weak force?*
- *Gravity? Is there a unified description of all forces?*
- *Why is $\text{mass}(W/Z/H) \ll \text{mass}(\text{Planck})$? (Hierarchy problem)*
- *Why is charge quantized?*
- *What is Dark Matter and Dark Energy? (and why Dark Energy now?)*
- *What was the Big Bang?*

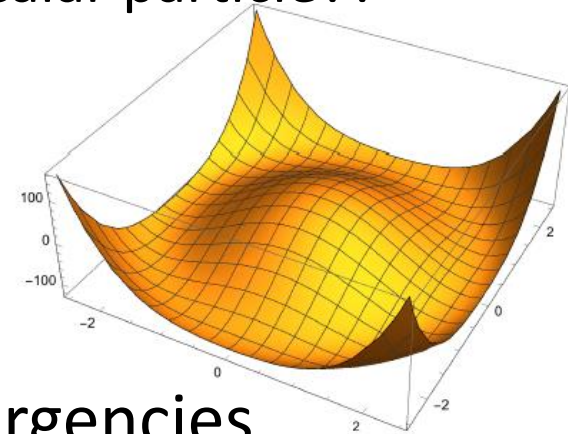
Unification of coupling constants?



Extrapolating the Standard Model coupling constants to higher energies

The Higgs discovery just adds to that list...

- What is it, really, a condensate in our Universe?
- Is it elementary?
 - If yes, why is there only 1 fundamental scalar particle??
- Why does it have mass² $\mu^2 < 0$?!
- Higgs mechanism gives quadratic divergencies
 - (see later)



Is the Standard Model really fundamental?

- Does not appear so (≈ 25 parameters?!)
- Evidence of selective processes:
 - For instance, no neutral colored fermions
 - $q_d = q_e / N(\text{colors}) \Rightarrow$ grand unification?
- **Fragile:** small changes in parameters \Rightarrow very different physics!
 - If $m_d < m_u$: all protons decay \Rightarrow no atoms
 - If $m_e > 4m_p - m_\alpha \Rightarrow$ Sun doesn't burn \Rightarrow no us
 - If $v \gg \text{TeV} \Rightarrow |m_n - m_p|$ large, rapid neutron decay \Rightarrow no chemistry nor life

The “Gauge Hierarchy Problem”

Discover of Higgs boson with mass < 1 TeV means the Standard Model is complete !

However, when computing radiative corrections to the bare Higgs mass a problem occurs:

The diagram shows two Feynman diagrams for Higgs mass corrections. The left diagram shows a top quark loop (t and t-bar) connected to two Higgs bosons (H). The right diagram shows a loop of gauge bosons (H, W, Z, gamma) connected to two Higgs bosons (H). A red arrow points from the diagrams to the equation below.

Higgs radiative corrections

$$m_H^2 = m_0^2 + \delta m_H^2 \quad \text{where:} \quad \delta m_H^2 \propto \int_0^\infty d^4k \frac{k^2 + m_f^2}{(k^2 + m_f^2)^2} + \dots \xrightarrow{\text{cut-off}} \int_0^{\Lambda_{\text{cut-off}}} (\dots) \propto \Lambda_{\text{cut-off}}^2$$

Integral quadratically divergent

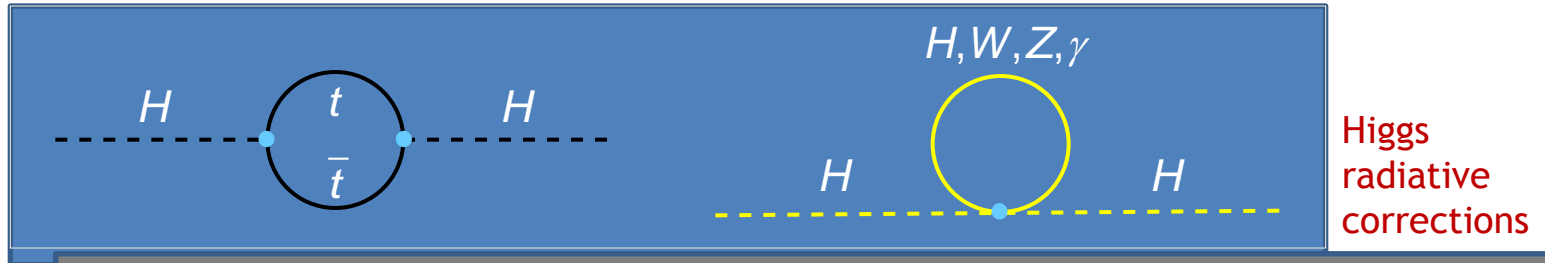
The cut-off sets the scale where new particles and physical laws must come in
 Above the EW scale we only know of two scales: GUT ($\sim 10^{16}$ GeV) and Planck ($\sim 10^{19}$ GeV)
 Such a cut-off would require an incredible amount of finetuning to keep m_H light

$$m_H^2 = (125 \text{ GeV})^2 = m_0^2 + C \cdot \Lambda_{\text{cut-off}}^2$$

The “Gauge Hierarchy Problem”

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$m_H^2 = m_0^2 + \delta m_H^2$ where:

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Integral quadratically divergent

Missing protection of scalar Higgs mass is related to **absence of a symmetry principle**. Setting $m_H = 0$ in SM Lagrangian, **does not restore any symmetry in the model**.

New physics models should address this. M_H should become a deviation from some exact symmetry, and is thus **intrinsically small** !

$$m_H^2 = (125 \text{ GeV})^2 = m_0^2 + C \cdot \Lambda_{\text{cut-off}}^2$$

Hunting for Answers

- ❖ Get more information
 - Measure particles and their interactions in detail
 - Precision measurements (e.g. LHCb)
 - Observe new particles or interactions
 - Search in new areas in “phase space”
- ❖ Find the underlying pattern(s)
 - Hypothesize, build models
 - Internally consistent? Consistent with data?
 - Suggestions on where to look

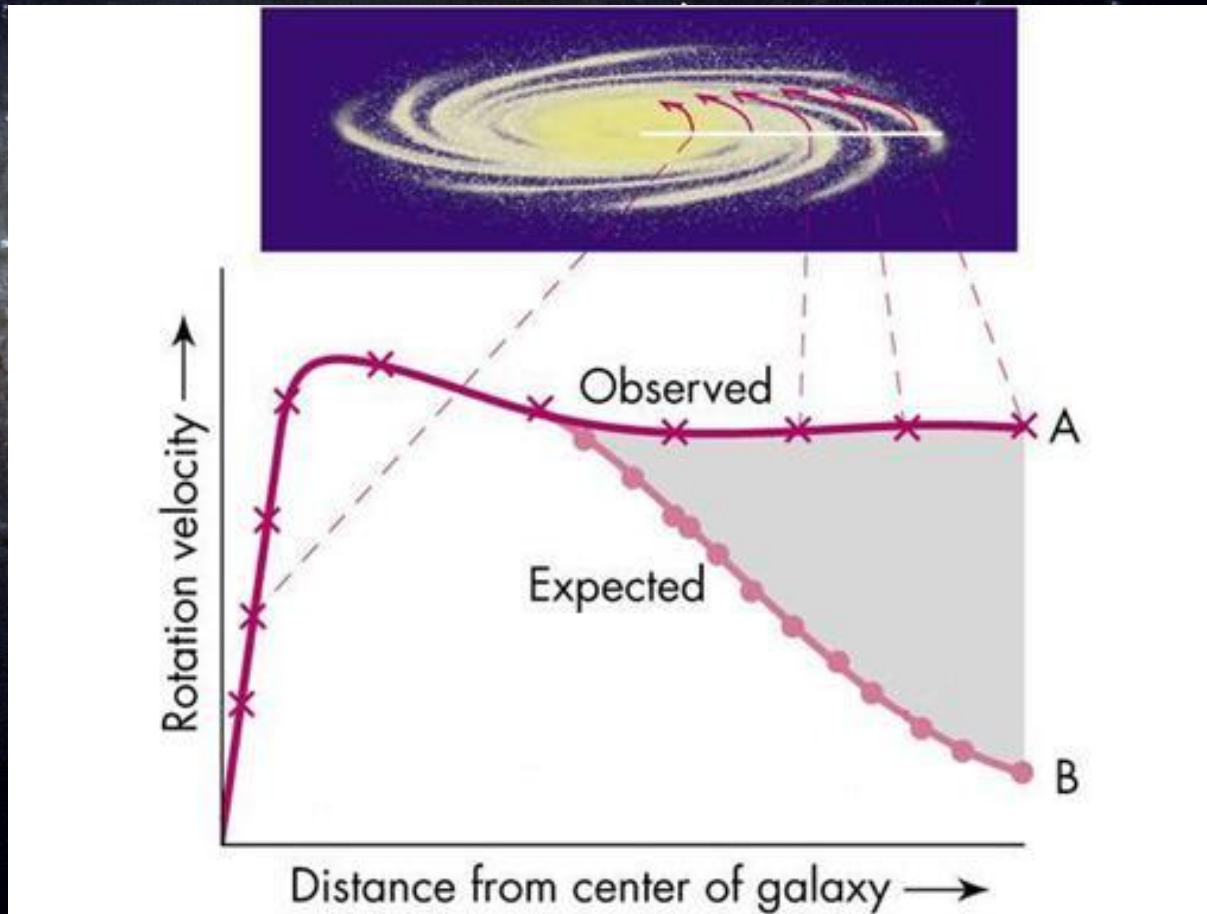
Experiment

Theory

Where to Start?

- ❖ BSM physics **must** couple to SM (weakly?), but is it
 - Resonant?
 - Does it have new massive particles decaying to electrons, muons, quarks, W, Z, ...?
 - “SM-like”?
 - Same but includes some new long-lived particles in the decay chain... (e.g. dark matter candidate)
 - No new “particles” in reach
 - Hidden or too heavy or... don't exist
 - Are there new interactions?

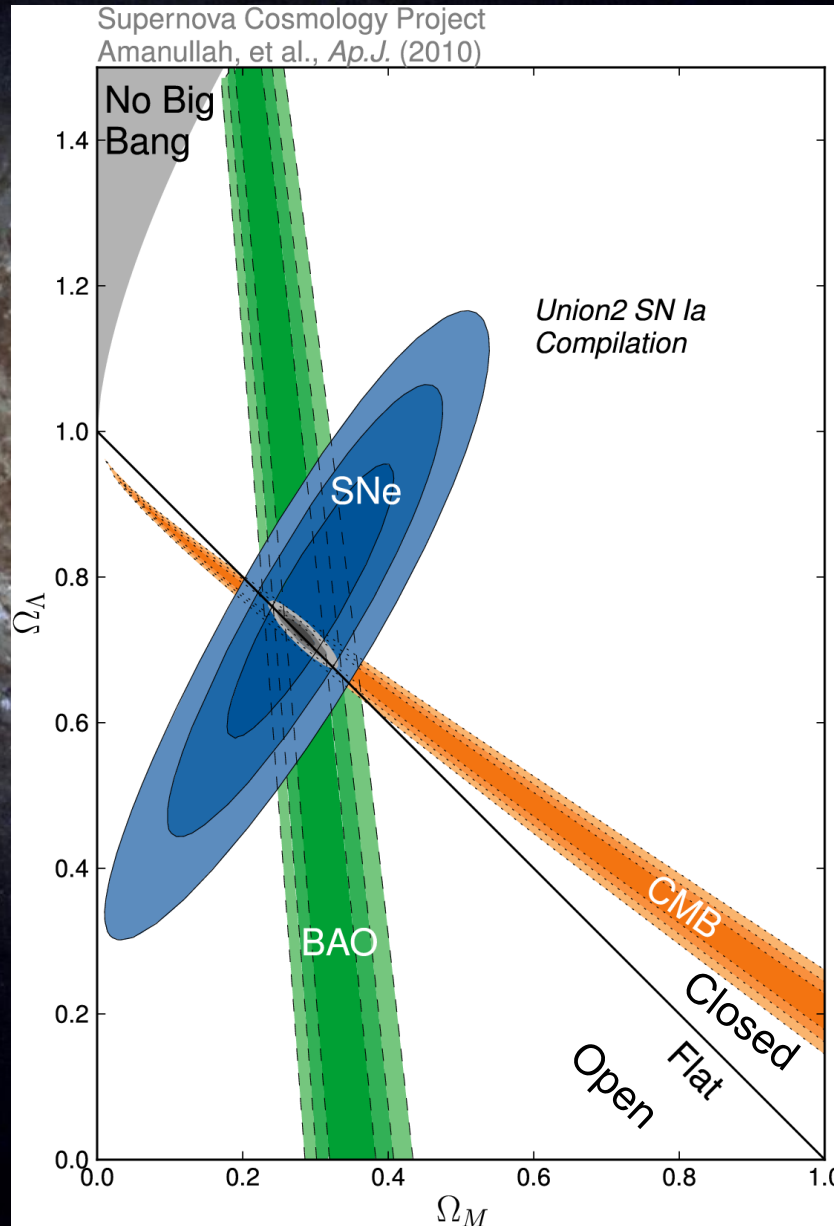
Galaxy rotation curves



**Standard Model only accounts for
~20% of the matter of the Universe!!!**

Supernovae data

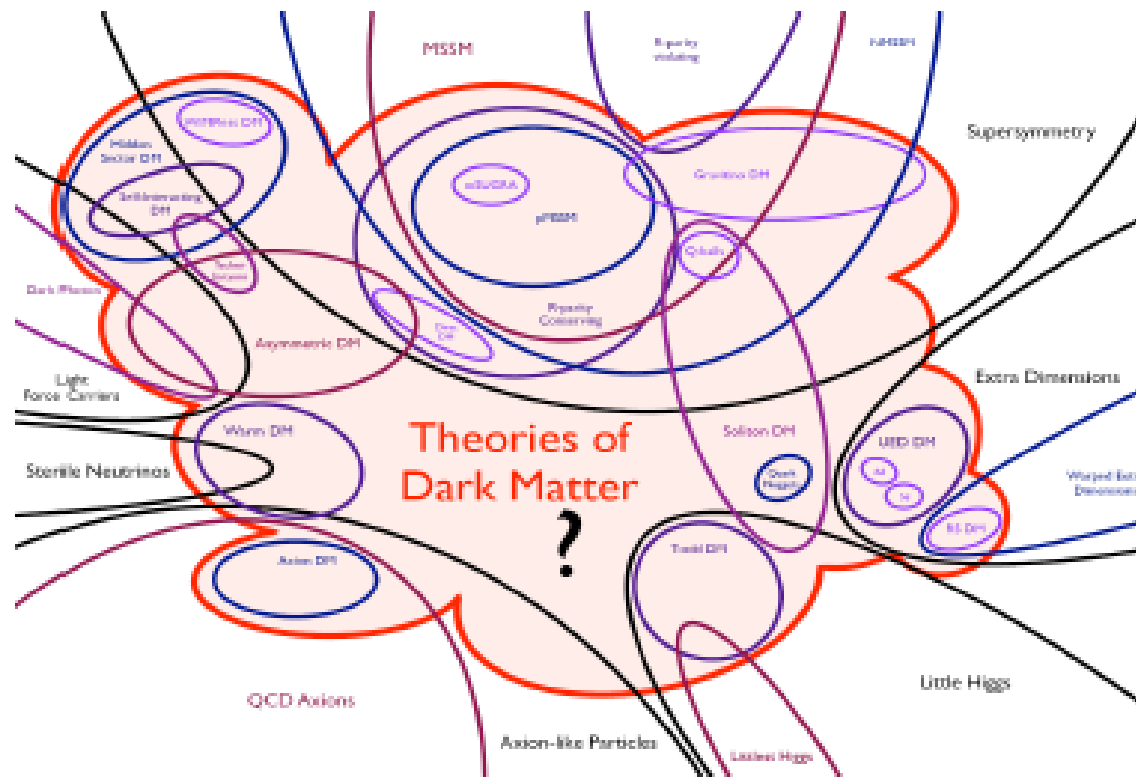
“Cosmological constant” term



Matter only
accounts for
~30% of the
Universe!

“Matter density” term

The energy scale(s) of new physics



T. Tait, DM@LHC '14

The prediction about the mass scale of DM comes with large error bars:

$$10^{-22} \text{ eV} < m_{DM} < 10^{20} \text{ GeV}$$

(ALPs)

(Wimpzillas, Q-balls)

Supersymmetry (SUSY)

Idea

New symmetry *fermions* \leftrightarrow *bosons*

This symmetry is the most general extension of Lorentz invariance

SUSY has: $N_{\text{dof}}(\text{bosons}) = N_{\text{dof}}(\text{fermions})$
 [cf. SM: $N_{\text{dof}}(\text{bosons}) \ll N_{\text{dof}}(\text{fermions})$]

| Spin 0 | Spin 1/2 | Spin 1 | Spin 3/2 | Spin 2 |
|----------|----------|--------|-----------|----------|
| sLeptons | Leptons | | Gravitino | Graviton |
| sQuarks | Quarks | | | |
| Higgs | Higgsino | | | |
| | Photino | Photon | | |
| | Zino | Z | | |
| | Wino | W | | |
| | Gluino | Gluon | | |

- To create *supermultiplets*, we need to add one *superpartner* to each SM particle
- Superpartners have opposite spin statistics but otherwise equal quantum numbers
- Need to introduce an additional Higgs doublet to the non-SUSY side \rightarrow 5 Higgs bosons

But where are these partners?!

Supersymmetry must be broken (if realized)

Particle spectrum (minimal!)

In reality the new states would mix

Several ideas of how the supersymmetry is broken – intimately connected with EWK symmetry breaking

| | | Spin 0 | Spin 1/2 | Spin 1 |
|--|---------------------|----------------------------------|--|----------------------|
| <div style="display: flex; flex-direction: column; align-items: center;"> <div style="border: 1px solid black; padding: 2px; margin-bottom: 2px;">SM</div> <div style="border: 1px solid black; padding: 2px;">SUSY</div> </div> | Eigenstates of mass | $\tilde{\ell}_1, \tilde{\ell}_2$ | ℓ | |
| | | \tilde{q}_1, \tilde{q}_2 | q | |
| | | h^0, H^0, A^0, H^\pm | $\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$ | |
| | | | $\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm$ | γ, Z^0, W^\pm |
| | | \tilde{g}_a | g_a | |

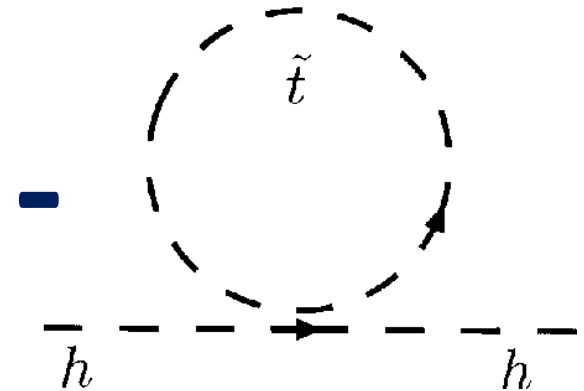
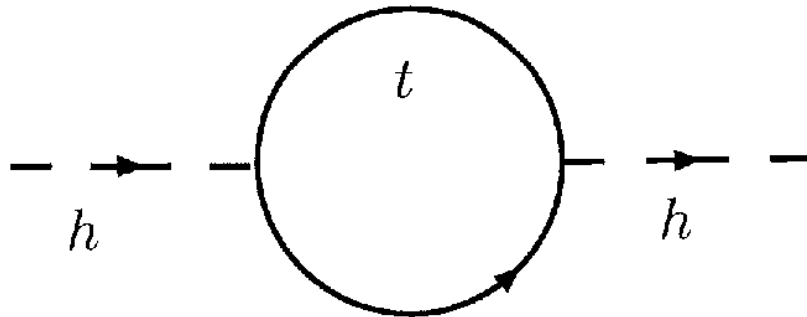
The gauge-mixed physical states that propagate in space and time and that can be observed.
 Neutralinos: mass eigenstates of photinos, zinos, neutral higgsinos
 Charginos : mass eigenstates of winos and charged higgsinos

Squark/slepton mixing proportional to SM partner masses
 → largest for 3rd gen.
 → can become lightest squarks / sleptons

Since we don't know the mechanism, have to introduce $\mathcal{O}(100)$ new parameters

SUSY and the hierarchy problem

If Supersymmetry not broken we would have perfect cancellation in the loops!



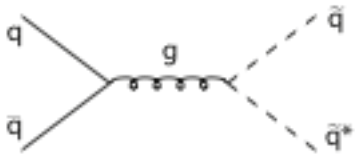
But as $m(\tilde{t}) \neq m(t)$ they do not quite cancel, instead just a suppression

This still gives a decent result if $|m(\text{fermion}) - m(\text{boson})| < \mathcal{O}(\text{TeV})$

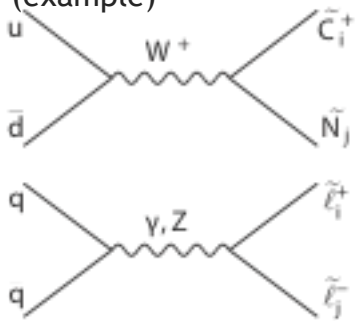
Once mass spectrum fixed, all cross sections predicted

Spin structure of SUSY spectrum: lower σ than other BSM models, harder to find !

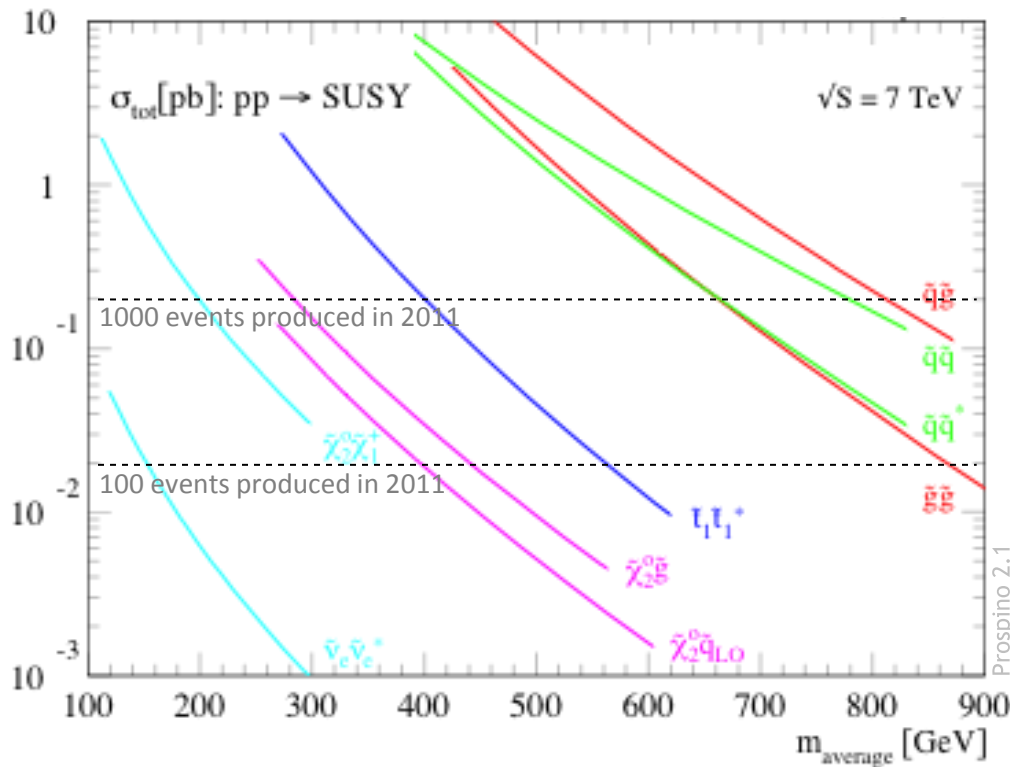
Direct squark pair production (example)



Direct gaugino/slepton pair production (example)

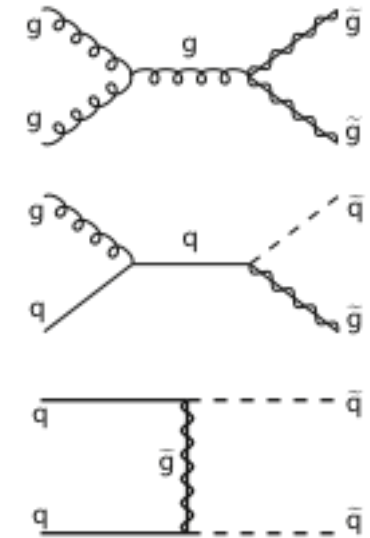


SUSY cross section versus sparticle mass

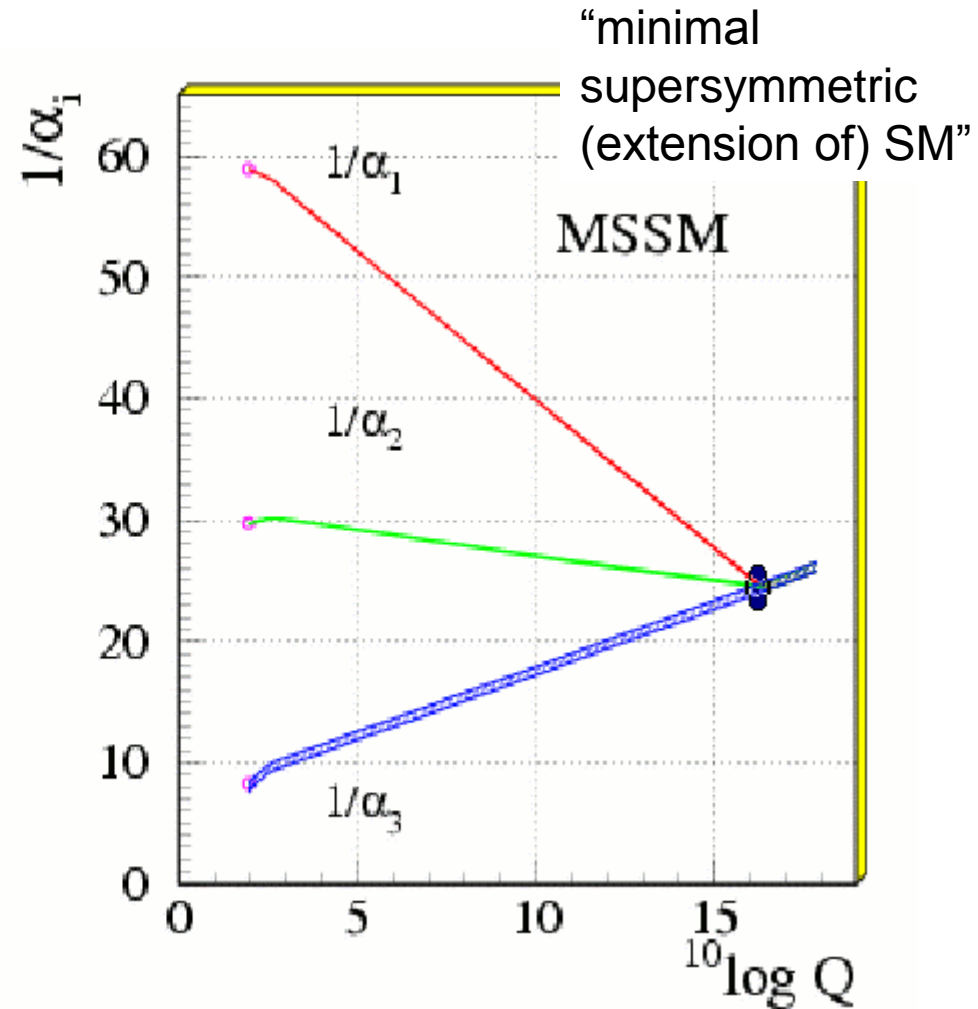
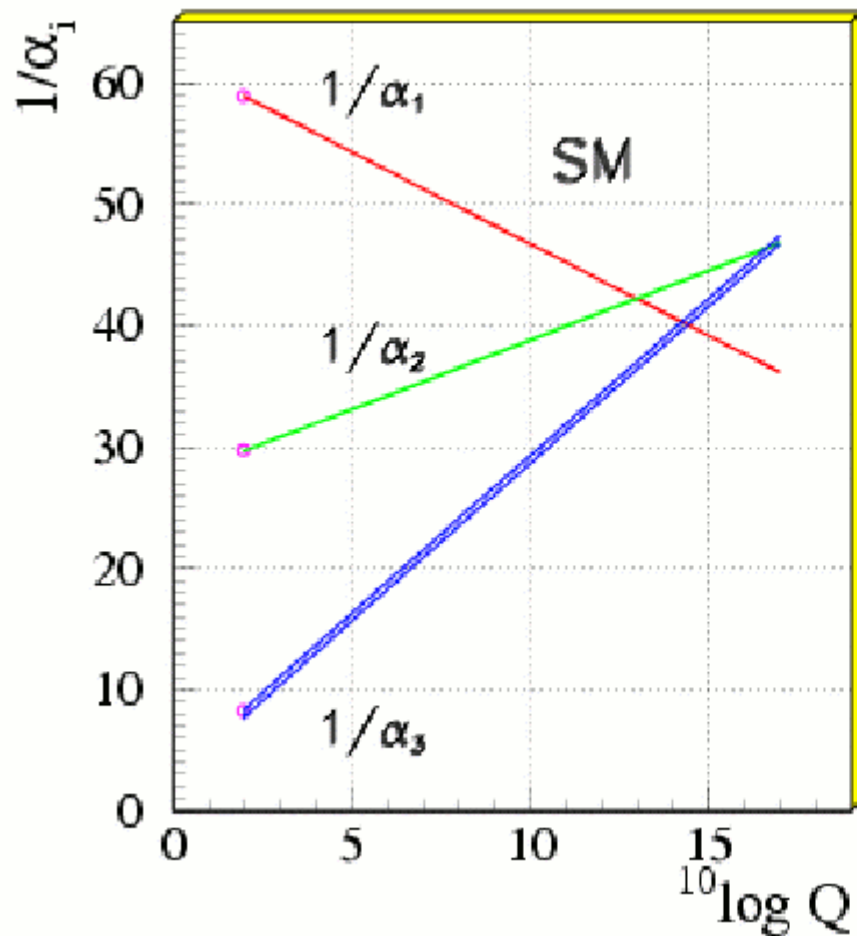


Prospino 2.1

Gluino & squark production (examples)



Unification of coupling constants with supersymmetry

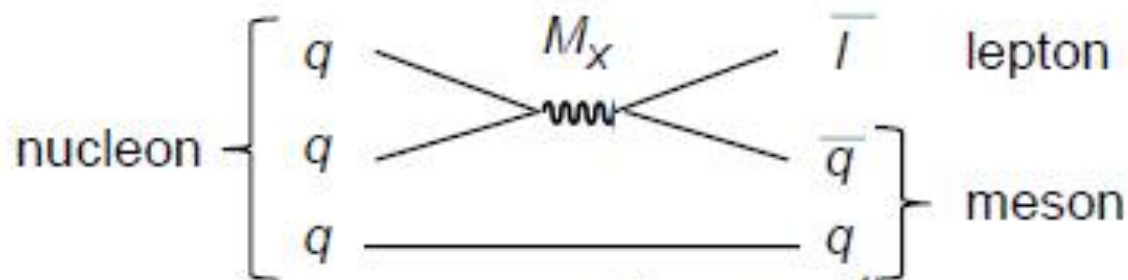


Proton Decay


(G. Giudice SSLP'15)

in GUT, matter is unstable

decay of proton mediated by new SU(5)/SO(10) gauge bosons



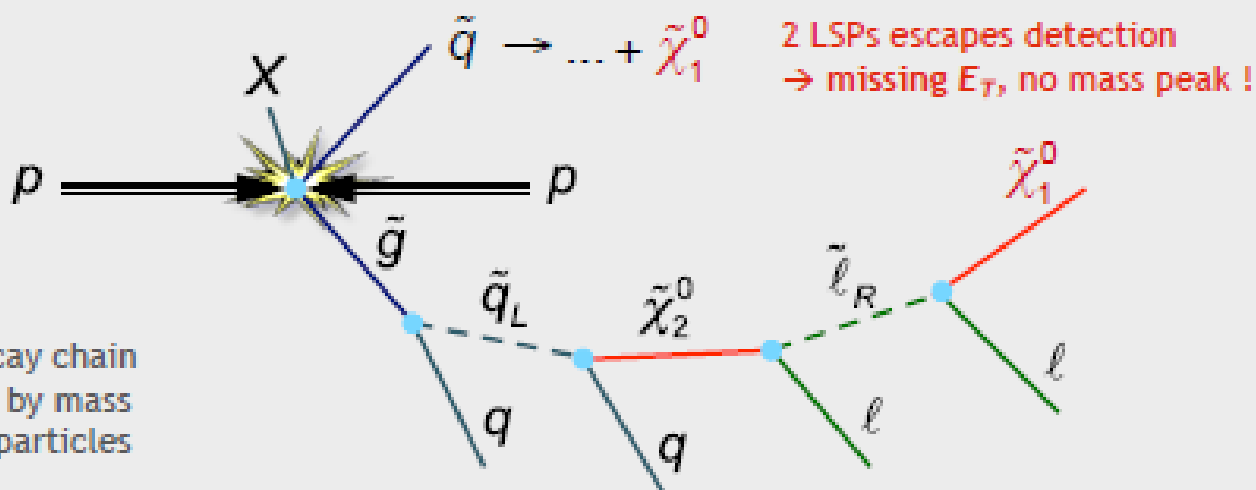
$$\text{GUT: } \tau_p(p \rightarrow e^+ \pi^0) = \left(\frac{M_X}{10^{15} \text{ GeV}} \right)^4 10^{31-32} \text{ yr}$$



$$\text{Exp: } \tau_p(p \rightarrow e^+ \pi^0) > 8.2 \times 10^{33} \text{ yr}$$

Characteristic SUSY Decay Cascades

- To avoid proton decay, a new conserved quantum number (R) is introduced, which forces a SUSY particle to decay in at least one other SUSY particle
- The lightest SUSY particle is thus stable (LSP), and must be neutral and colourless \rightarrow WIMP (dark matter candidate)
- Typical **LSP is spin- $\frac{1}{2}$ neutralino**. It could also be a gravitino
- With R parity: SUSY production in pairs only \rightarrow requires energy $2 \times$ SUSY mass !

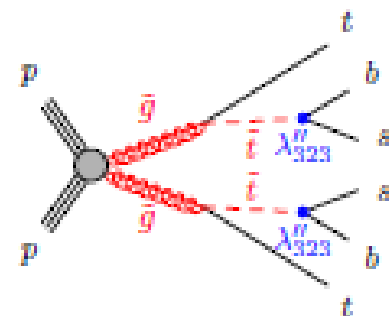
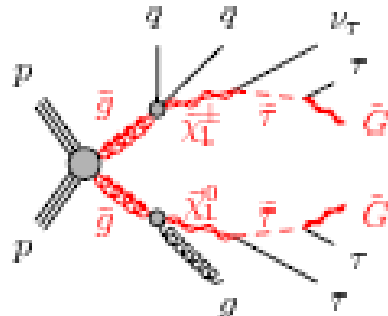
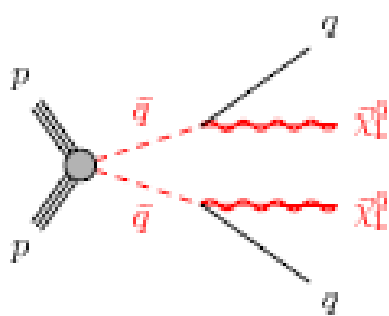


"Typical" SUSY decay chain at the LHC, driven by mass hierarchy of SUSY particles

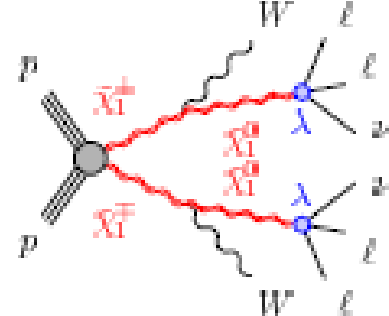
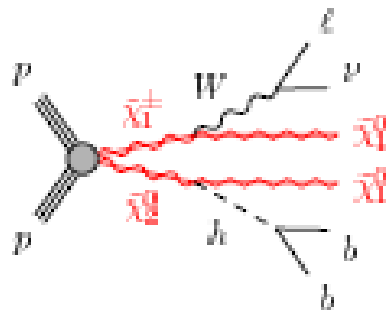
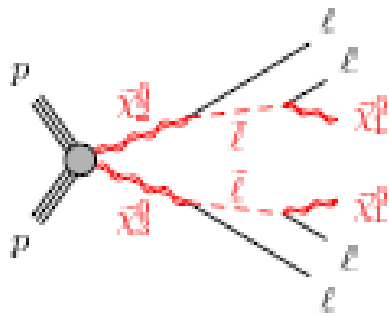
Canonical SUSY

❖ Wide range of signatures

- Strong production... (large cross-section)



- ... or weak production



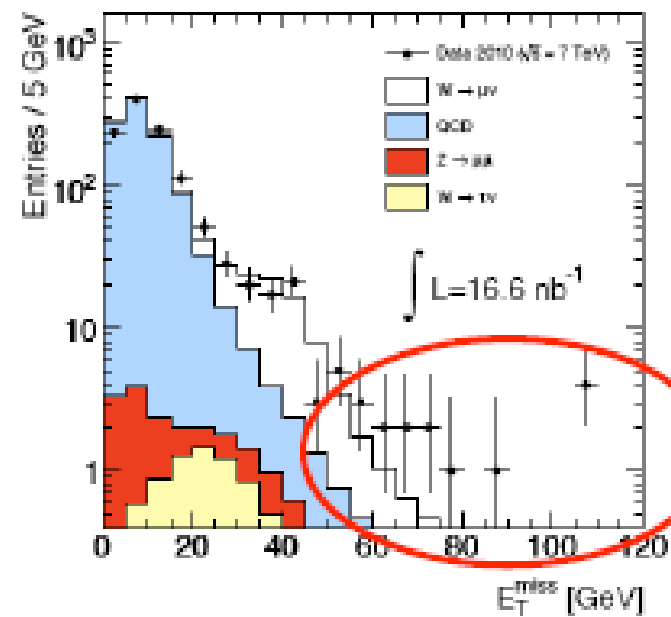
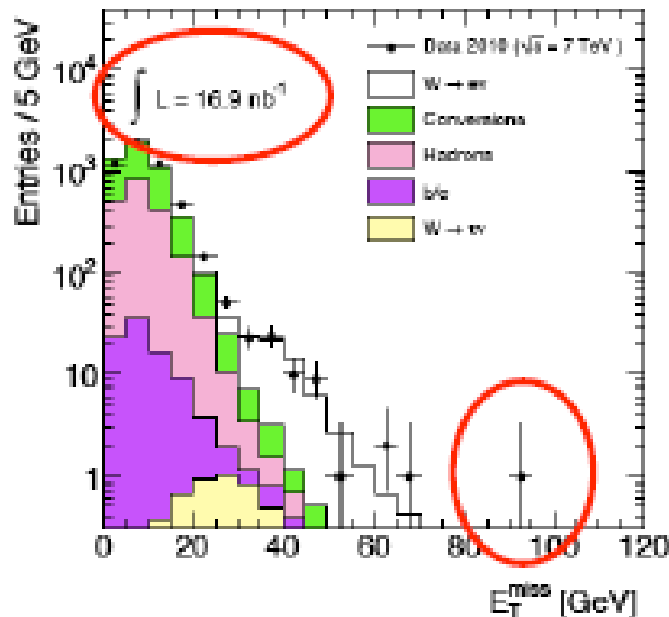
RPV

Missing ET

❖ “Evil” variable: - Σ (everything else)

- Need to understand “everything else”
- Good benchmark: leptonic W boson decays

Early 2010



❖ Analyses using MET are particularly sensitive

- Requires the full calorimeter to behave, and calorimeter is generally the most sensitive subdetector (analog, ~ 16 bits)
- Easy: basic DQ (high voltage trip, etc.)
- Hard: low frequency
- Can't spot a 10^{-5} Hz (once a day) effect online or in first pass DQ
- But can be biggest part of dataset after cuts!

