

Search for charginos

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Mass eigenstates of the supersymmetric partners to the Higgs and Gauge bosons:

$$[\tilde{W}^\pm, \tilde{H}^\pm] \Rightarrow \tilde{\chi}_i^\pm \quad i=1,2$$
$$[\tilde{\gamma}, \tilde{Z}^0, \tilde{H}_{1,2}^0] \Rightarrow \tilde{\chi}_i^0 \quad i=1,..,4$$

- R-parity conservation

→ Stable **L**ightest **S**upersymmetric **P**article.

→ Pair production of the SUSY particles.

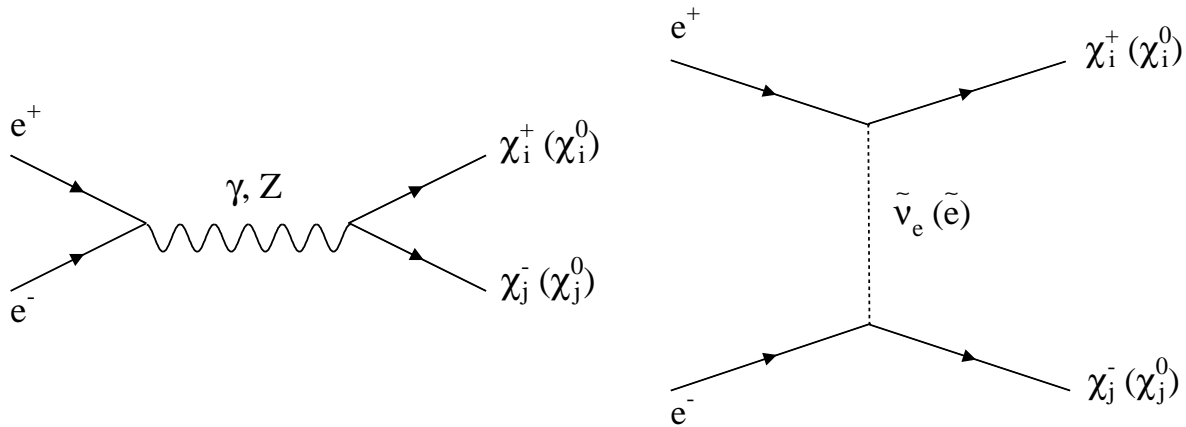
- GUT

→ 8 parameters, $[m_0, m_{\frac{1}{2}}, m_A, \mu, \tan(\beta), A_i]$.

- SUGRA

→ $\tilde{\chi}_1^0$ ($\tilde{\nu}$) is the LSP.

When the Chargino is not pure Higgsino the t-channel diagram contributes with negative interference for low $m_{\tilde{\nu}}$.

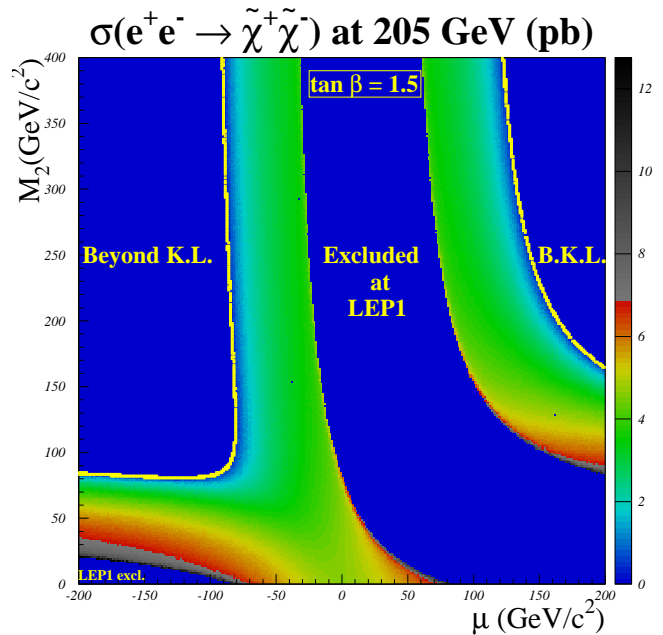
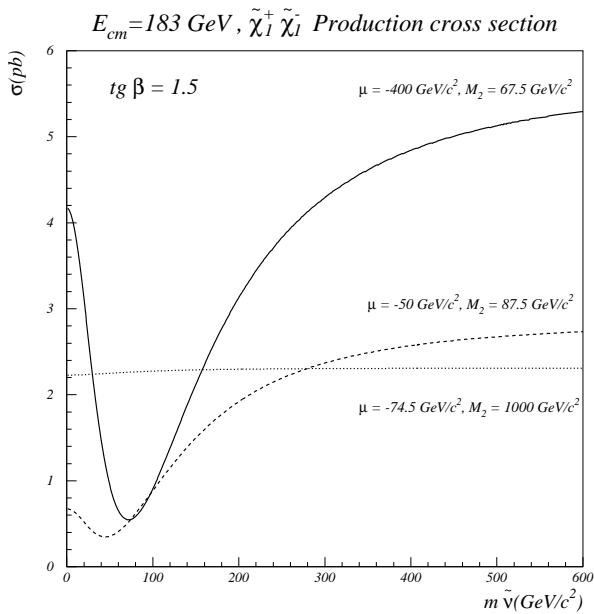


High $m_{\tilde{\nu}}$
(Constrained by m_0 , $m_0 = 1$ TeV)



$$M_{\tilde{\chi}_i^\pm} = f(\mu, \tan(\beta), m_{\frac{1}{2}})$$

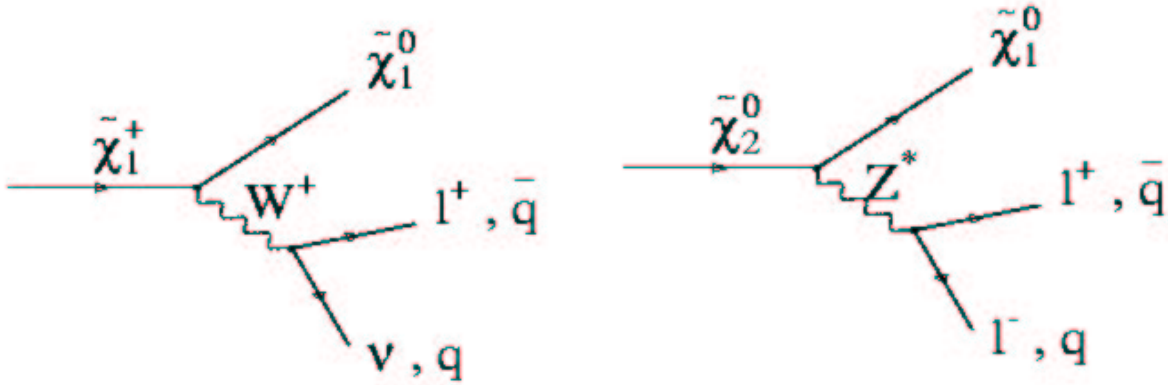
$$\sigma(e^+e^- \rightarrow \tilde{\chi}_i^+ \tilde{\chi}_j^-) = f(\mu, \tan(\beta), m_{\frac{1}{2}}, E_{cm})$$



The signal shape depends on:

$$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0} = \Delta M$$

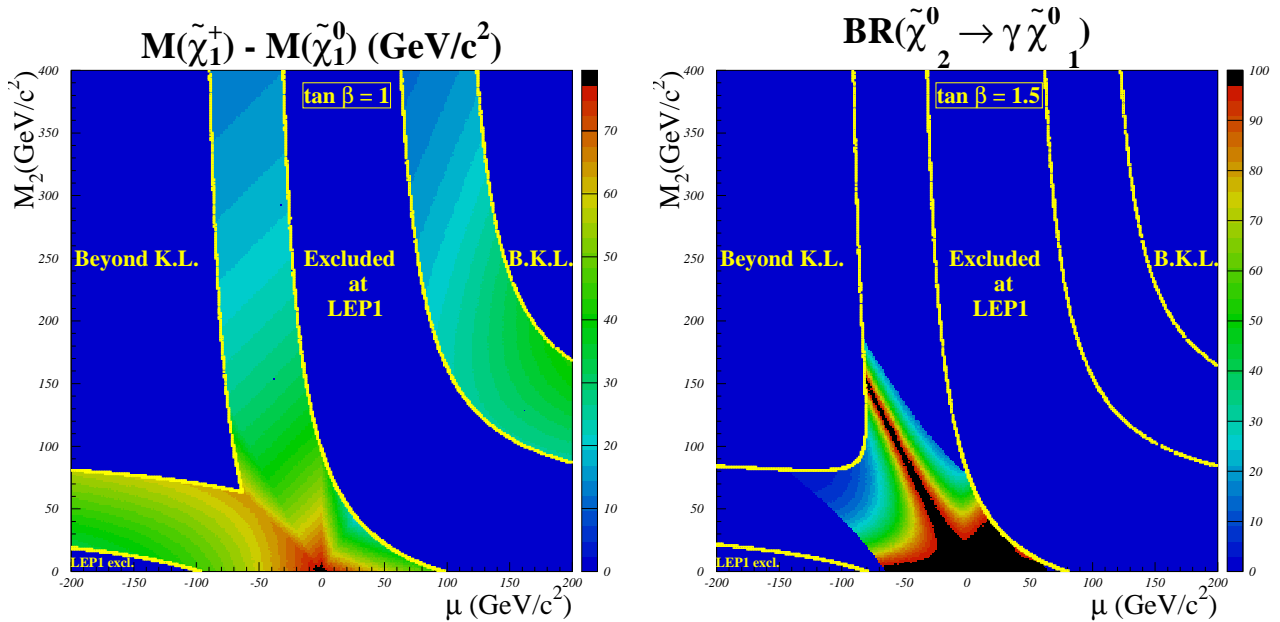
$$(m_{\tilde{l}, \tilde{q}, H^\pm, 0, A} > M_{\tilde{\chi}_1^\pm})$$



And if $\tilde{\chi}_2^0$ can be produced, causing a cascade decay:

$$\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_2^0 f_1 \bar{f}'_1 \rightarrow \tilde{\chi}_1^0 f_1 \bar{f}'_1 + f_2 \bar{f}_2$$

$$\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_2^0 f_1 \bar{f}'_1 \rightarrow \tilde{\chi}_1^0 f_1 \bar{f}'_1 + \gamma$$

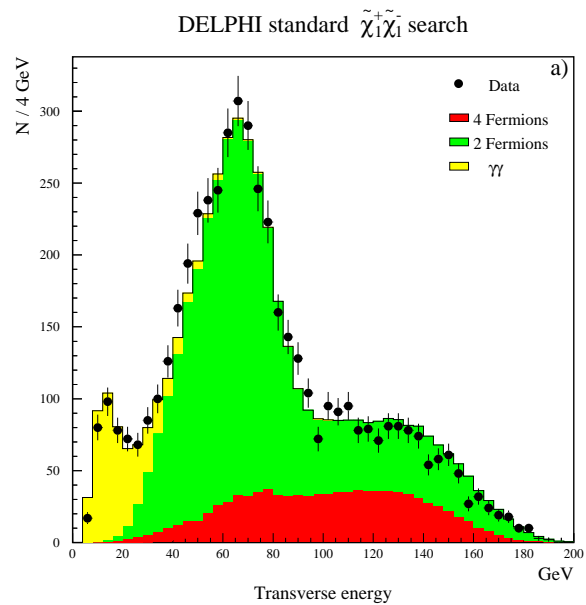
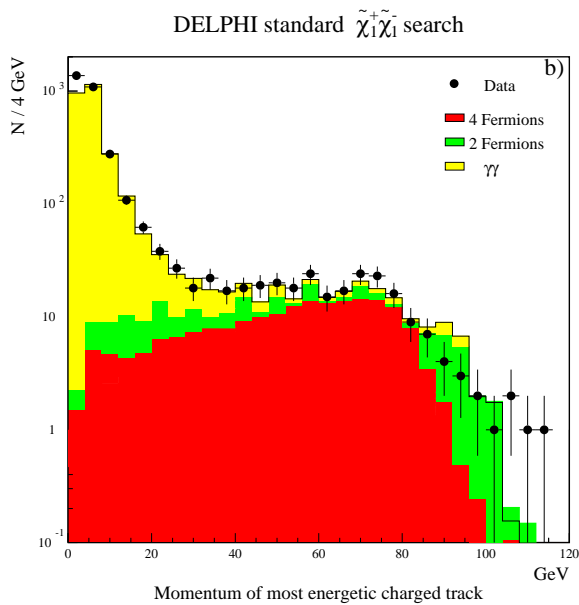


Main background



Low $\Delta M \rightarrow \gamma\gamma$
High $\Delta M \rightarrow 4/2$ fermions (WW)

$$E_{cm} = 206 \text{ GeV}$$



No isolated Photon
Number of tracks ≤ 5

No isolated Photon or lepton
Number of tracks > 5

- Radiative topology: Important for high ΔM
(WW suppression)

At least 1 isolated photon

($E_\gamma > 5 \text{ GeV}$; $15^\circ < \theta_\gamma < 165^\circ$; $\alpha_\gamma^{iso} > 15^\circ$)

- ll topology:

No isolated photons

Number of tracks ≤ 5

- j ℓ topology:

No isolated photons

Number of tracks > 5

At least 1 isolated lepton ($P_\ell > 3 \text{ GeV}/c$; $\alpha_\ell^{iso} > 20^\circ$)

- jj topology:

No isolated photons

Number of tracks > 5

No isolated leptons

ΔM REGIONS

$3 \leq \Delta M < 5$	GeV/c^2
$5 \leq \Delta M < 10$	GeV/c^2
$10 \leq \Delta M < 25$	GeV/c^2
$25 \leq \Delta M < 35$	GeV/c^2
$35 \leq \Delta M < 50$	GeV/c^2
$50 \leq \Delta M$	GeV/c^2

EVENTS



for a given topology

Topological cuts
(the same for all ΔM)



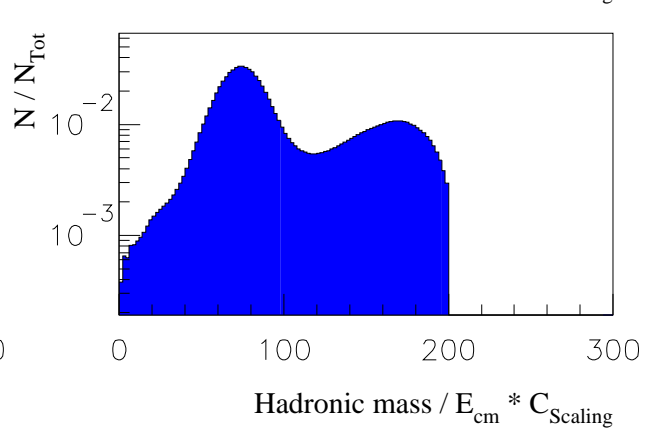
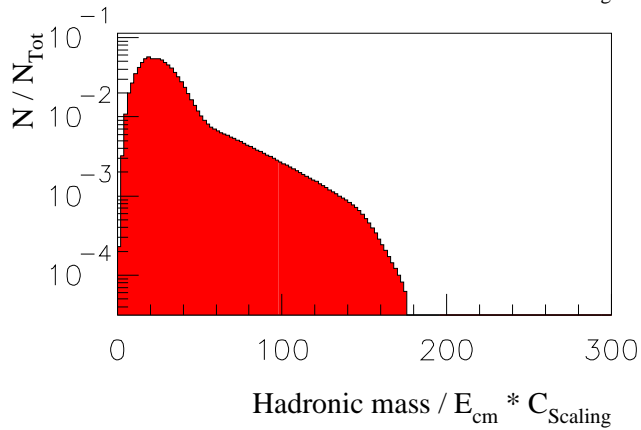
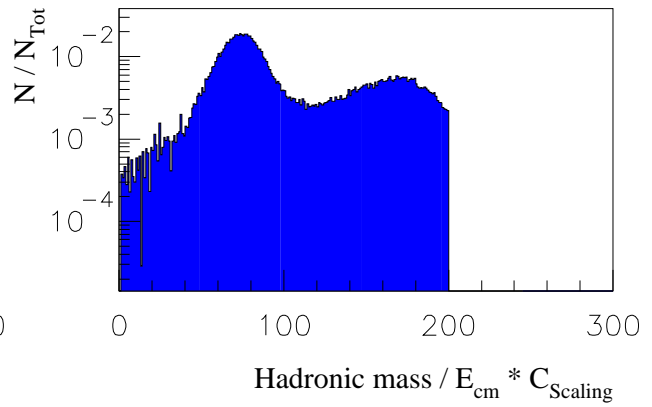
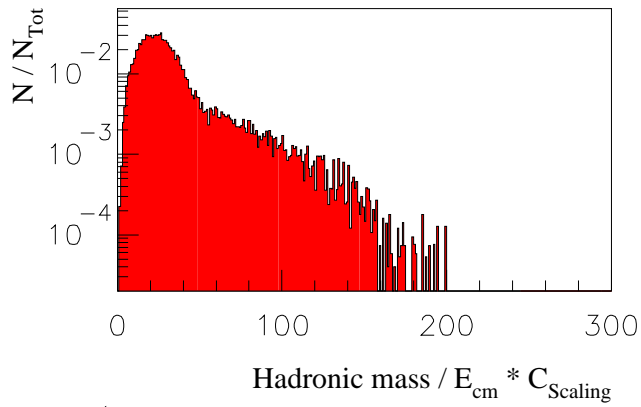
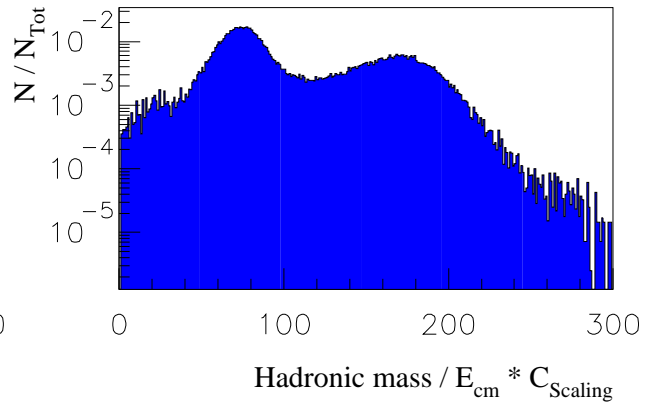
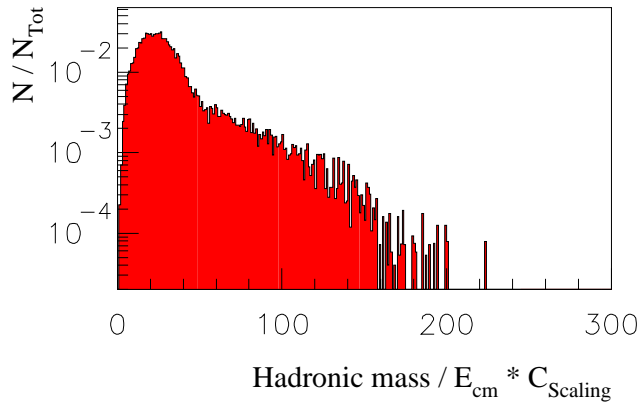
Preselection cuts
(different for each ΔM)



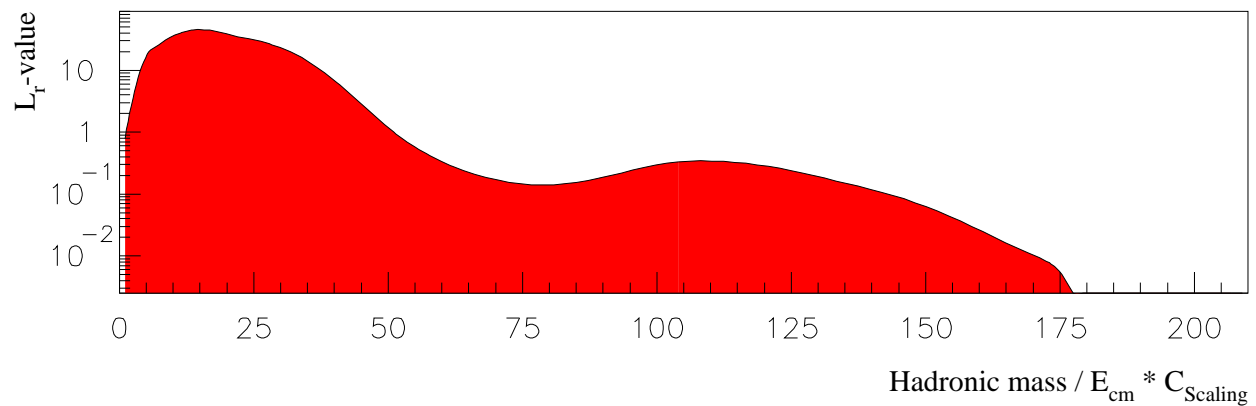
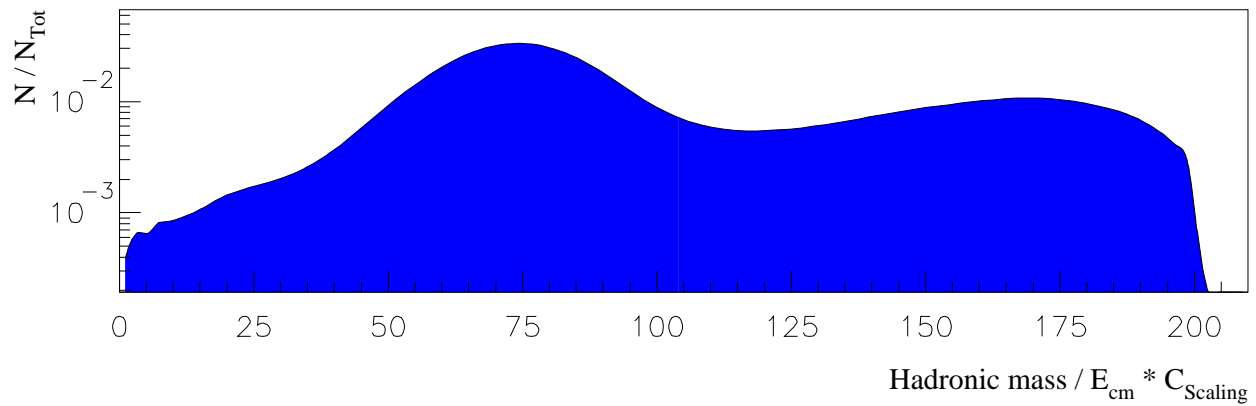
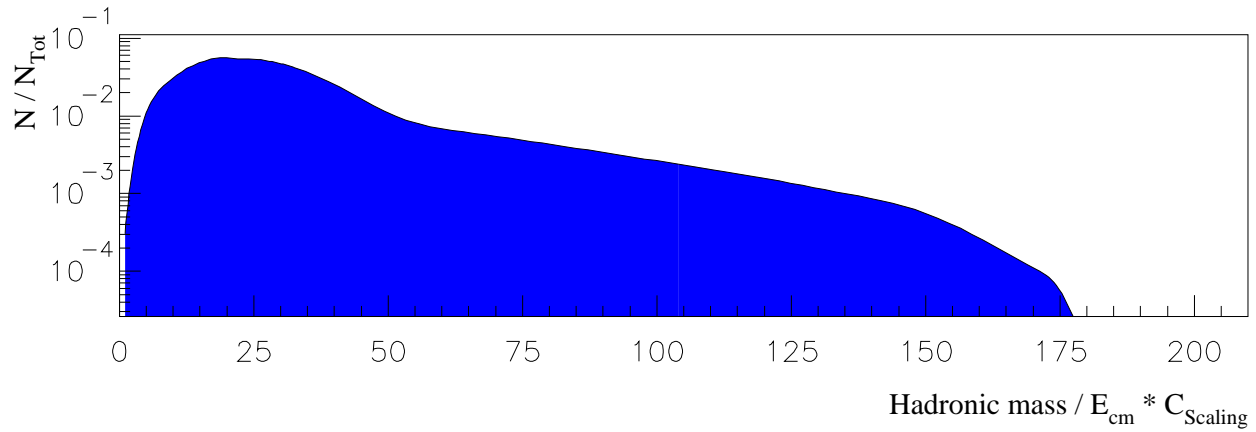
Likelihood cut

$$(\mathcal{L}_{\mathcal{R}}(x_1, \dots, x_n) = \frac{\prod_{i=1}^n P_S^i(x_i)}{\prod_{i=1}^n P_B^i(x_i)})$$

- Cut away events with $P_S() = 0$
- Triangular filter

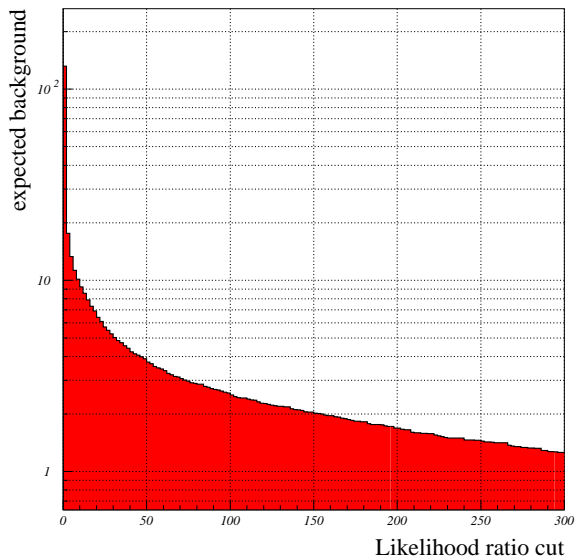


$$\mathcal{L}_{\mathcal{R}}(M_{Hadronic}) = \frac{P_S(M_{Hadronic})}{P_B(M_{Hadronic})}$$

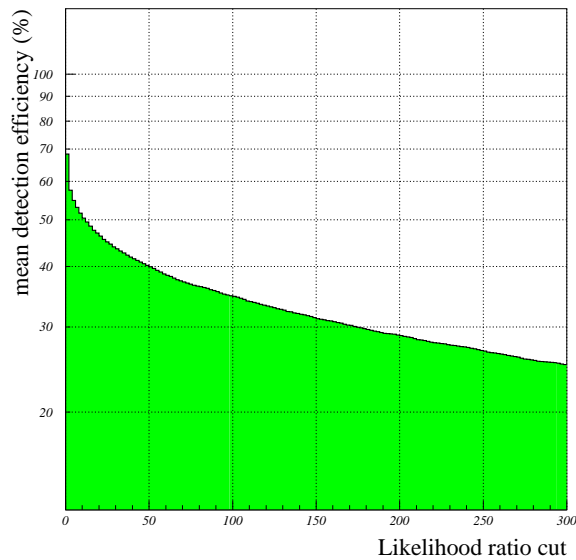


$$\sigma_{exclusion} = \frac{N_{95}(N_B(\mathcal{L}_{\mathcal{R}_{cut}}))}{\epsilon_S(\mathcal{L}_{\mathcal{R}_{cut}}) \times L}.$$

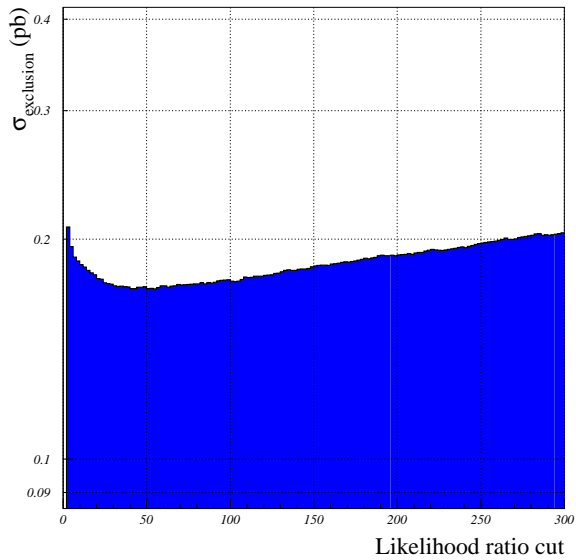
gg topology - ΔM region 2 - $E_{cm} = 200$ GeV



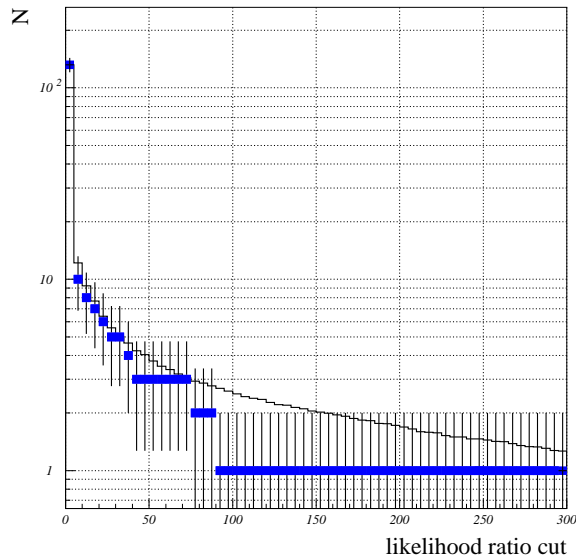
gg topology - ΔM region 2 - $E_{cm} = 200$ GeV



gg topology - ΔM region 2 - $E_{cm} = 200$ GeV



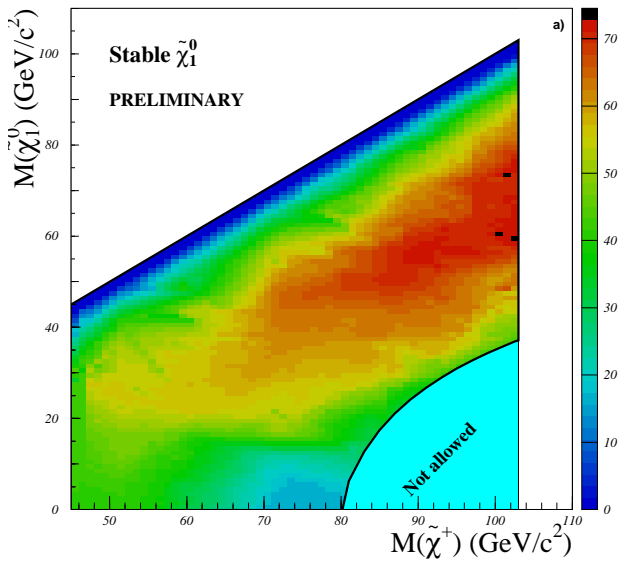
gg topology - ΔM region 2 - $E_{cm} = 200$ GeV



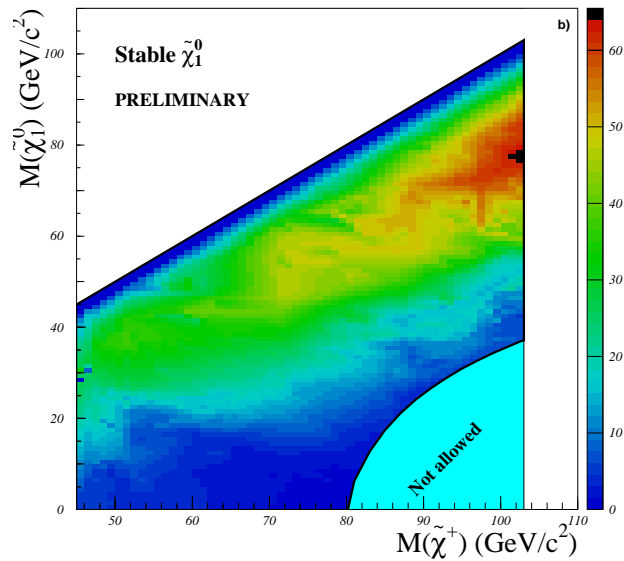
Topology:	$j\ell$	ll	jj	rad	TOTAL
$3 \leq \Delta M < 5 \text{ GeV}/c^2$					
N data:	1	54	1	6	62
N backg:	3.3 $^{+1.90}_{-0.64}$	54.8 $^{+4.67}_{-3.35}$	2.5 $^{+1.87}_{-0.56}$	4.2 $^{+1.84}_{-0.43}$	64.9 $^{+5.63}_{-3.47}$
$5 \leq \Delta M < 10 \text{ GeV}/c^2$					
N data:	1	16	4	6	27
N backg:	3.3 $^{+1.90}_{-0.64}$	10.5 $^{+2.43}_{-1.12}$	6.2 $^{+2.13}_{-0.91}$	4.2 $^{+1.84}_{-0.43}$	24.3 $^{+4.11}_{-1.69}$
$10 \leq \Delta M < 25 \text{ GeV}/c^2$					
N data:	1	12	8	6	27
N backg:	3.3 $^{+1.90}_{-0.64}$	11.3 $^{+2.18}_{-0.81}$	10.7 $^{+2.10}_{-1.05}$	4.2 $^{+1.84}_{-0.43}$	29.6 $^{+4.05}_{-1.58}$
$25 \leq \Delta M < 35 \text{ GeV}/c^2$					
N data:	2	26	7	6	41
N backg:	2.7 $^{+1.77}_{-0.39}$	21.2 $^{+2.39}_{-1.08}$	5.7 $^{+1.71}_{-0.38}$	4.2 $^{+1.84}_{-0.43}$	33.9 $^{+3.82}_{-1.23}$
$35 \leq \Delta M < 50 \text{ GeV}/c^2$					
N data:	2	30	16	14	62
N backg:	2.1 $^{+1.56}_{-0.12}$	36.4 $^{+2.52}_{-1.23}$	14.6 $^{+1.69}_{-0.43}$	13.5 $^{+1.84}_{-0.56}$	66.7 $^{+3.87}_{-1.46}$
$50 \text{ GeV}/c^2 \leq \Delta M$					
N data:	5	44	26	14	89
N backg:	7.9 $^{+1.74}_{-0.45}$	47.7 $^{+2.47}_{-1.21}$	25.0 $^{+1.73}_{-0.50}$	13.5 $^{+1.84}_{-0.56}$	94.2 $^{+3.97}_{-1.51}$
TOTAL (logical .OR. between different ΔM windows)					
N data:	7	101	37	15	160
N backg:	11.6 $^{+2.10}_{-0.82}$	105.6 $^{+4.91}_{-3.62}$	36.8 $^{+2.55}_{-1.59}$	15.1 $^{+1.98}_{-0.78}$	169.3 $^{+6.24}_{-4.19}$

Signal efficiencies at $E_{cm} = 206$ GeV

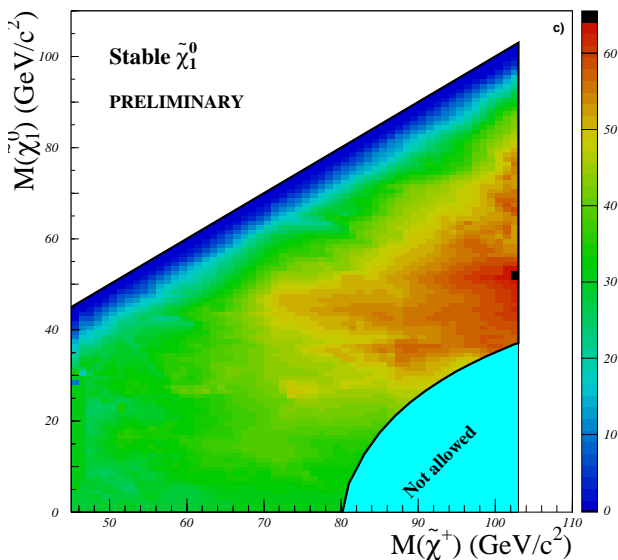
$e^+e^- \rightarrow \tilde{\chi}^+\tilde{\chi}^- \rightarrow jjl$ detection efficiency



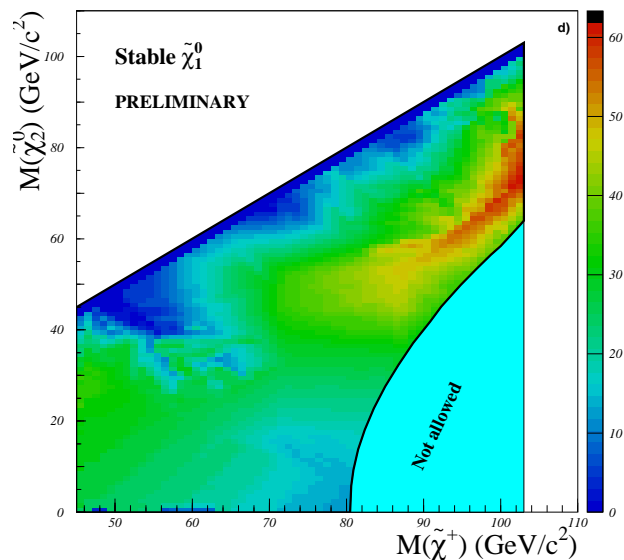
$e^+e^- \rightarrow \tilde{\chi}^+\tilde{\chi}^- \rightarrow$ jets detection efficiency



$e^+e^- \rightarrow \tilde{\chi}^+\tilde{\chi}^- \rightarrow ll$ detection efficiency



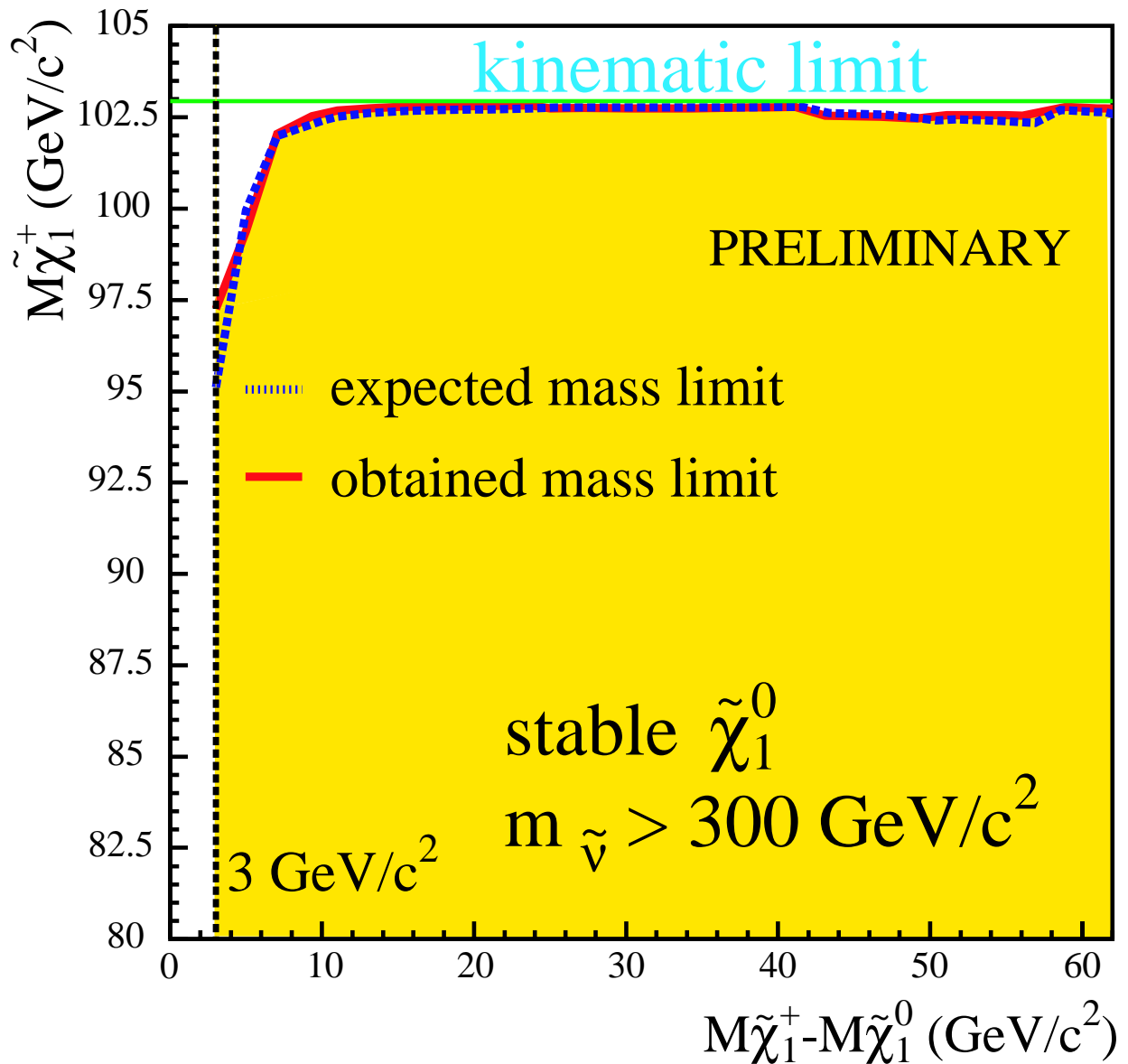
$e^+e^- \rightarrow \tilde{\chi}^+\tilde{\chi}^- \rightarrow \gamma\gamma X$ detection efficiency



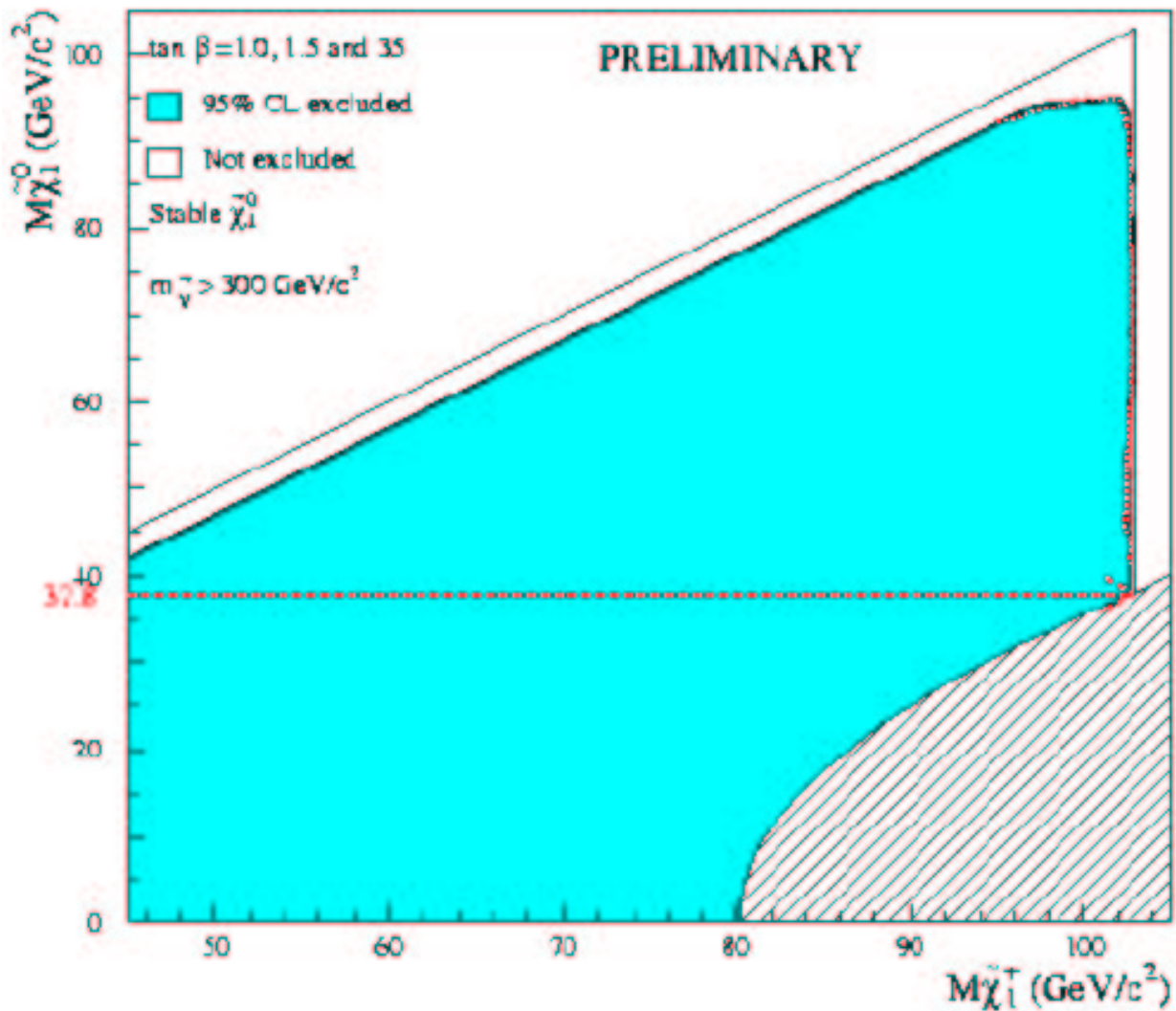
No Excess



Limits were set at 95 C.L.



DELPHI $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ mass limits



Implied limit on lightest neutralino mass

$$M_{\tilde{\chi}_1^0} = 37.8 \text{ GeV}/c^2 \text{ for } m_{\tilde{\nu}} > 300 \text{ GeV}/c^2$$

Present limits (sent to Moriond)

Case	$m_{\tilde{\nu}}$ (GeV/c ²)	$M_{\chi^{\pm}}^{min}$ (GeV/c ²)	σ^{max} (pb)
Stable neutralino $E_{cm} = 191.6-205.9$ GeV			
$\Delta M_+ > 10$ GeV/c ²	> 300	102.5	0.21
$\Delta M_+ = 5$ GeV/c ²	> 300	99.4	0.81
$\Delta M_+ = 3$ GeV/c ²	$> M_{\chi^{\pm}}$	97.3	1.09

- Until June (Summer conferences)

→ Redo the analysis while waiting for the final reprocessing.

New PDF construction

Lr-optimisation at 206 GeV

Including S1 data

Split of the 2000 data into four windows for the limit computation

Expected final limits

$$\Delta M > 10 \text{ GeV}/c^2 \Rightarrow 103$$

$$\Delta M = 5 \text{ GeV}/c^2 \Rightarrow 101$$

$$\Delta M = 3 \text{ GeV}/c^2 \Rightarrow 95$$

→ The effect of the interpolation of the signal efficiency on the systematics

- After June

→ Systematical errors

→ DELPHI note